

**Study of the energy storage process in the
magnetosphere using a novel nonlinear
frequency analysis**

**Zi-Qiang Lang, Yunpeng Zhu,
Richard Boynton, Michael Balikhin**

Department of Automatic Control and Systems
Engineering
University of Sheffield, UK

OUTLINE

Energy storage process in the magnetosphere

The Nonlinear AutoRegressive with eXogenous (NARX) input model of a magnetosphere system

A novel analysis of the energy storage process of the magnetosphere system

Revelation of energy storage phenomena with the magnetosphere system from satellite data analysis and possible significance

Conclusions

Energy storage process in the magnetosphere

V: Solar Wind Speed

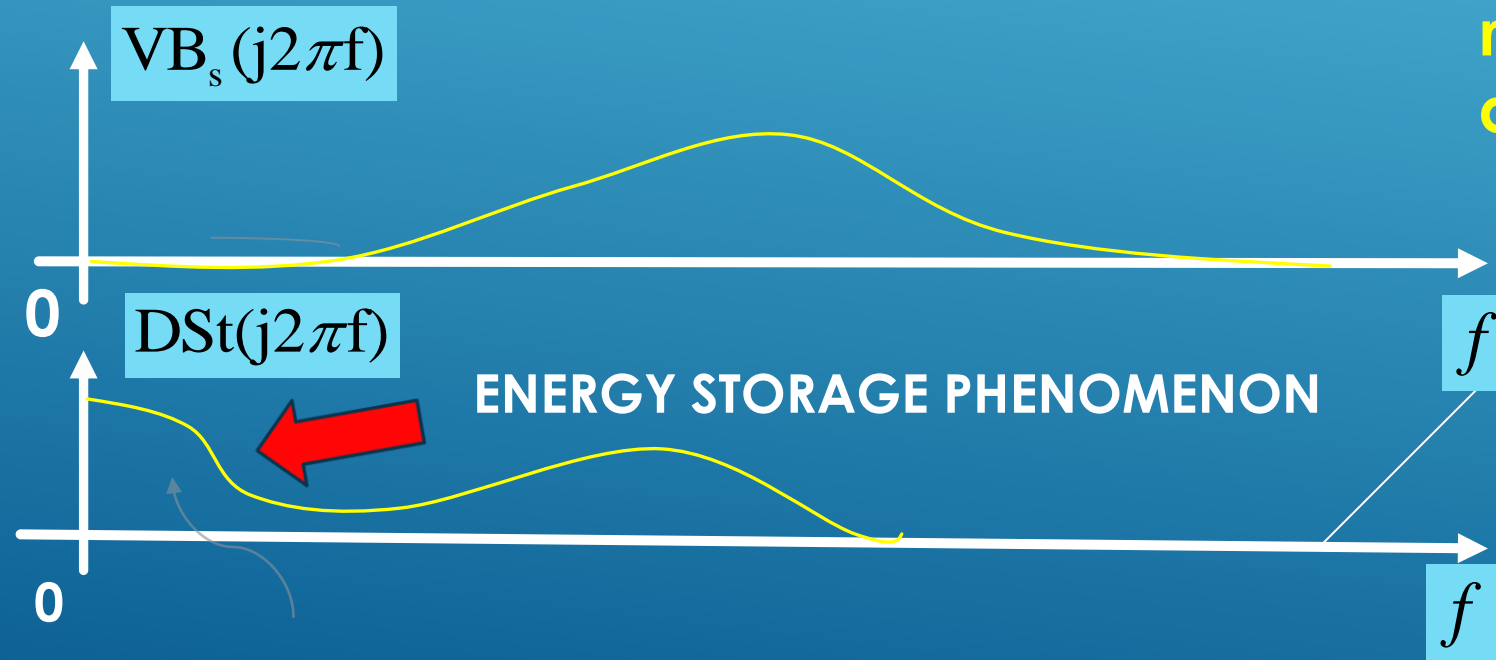
VB_s (INPUT)



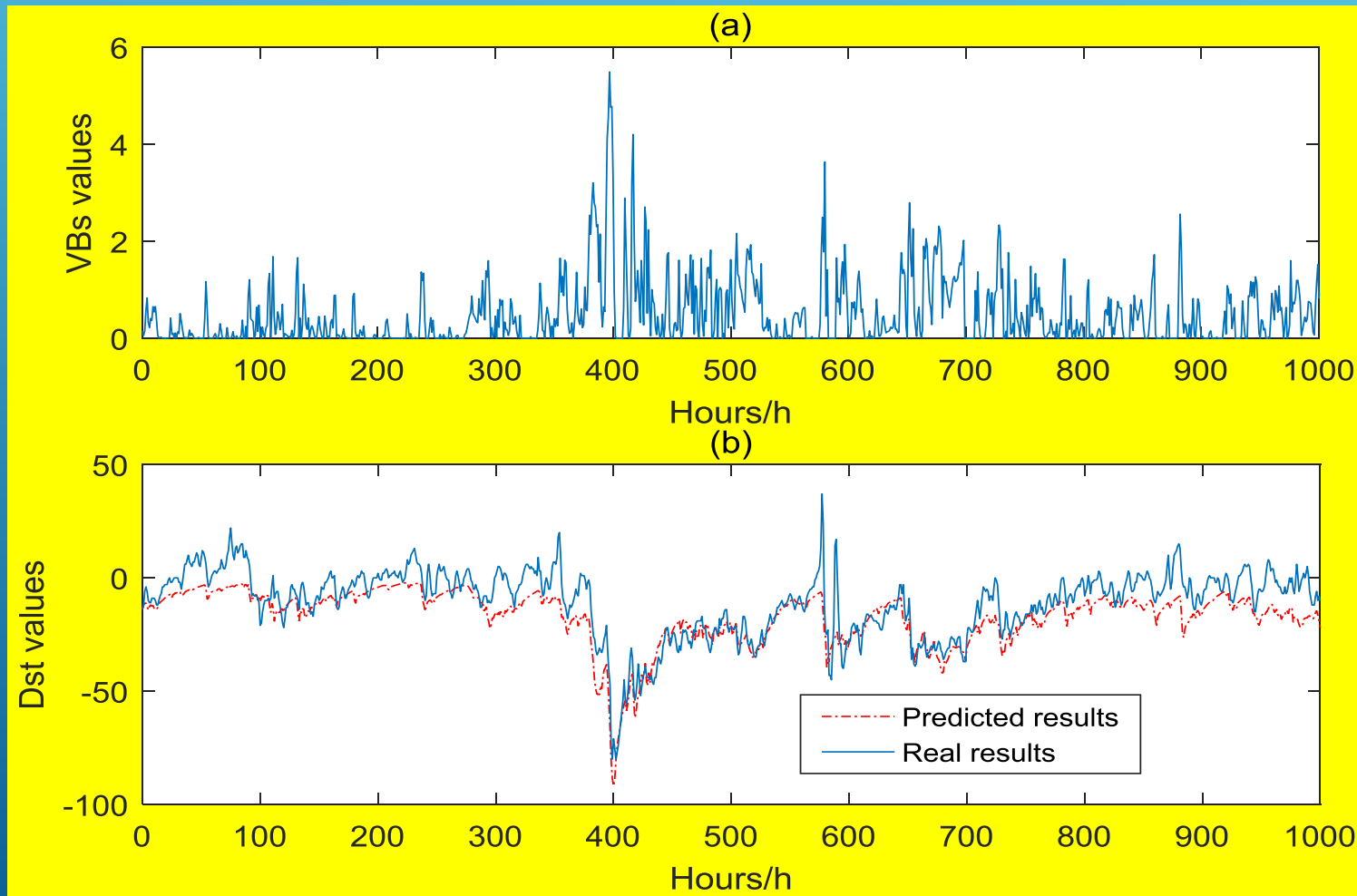
Dst (Output)

Index
quantifying
geomagnetic
storm
related
activity

B_s: Southen
Component
of Solar Wind
Megnet Field



The Nonlinear AutoRegressive with eXogenous (NARX) input model of a magnetosphere system



$$y(k) = 0.953y(k-1) - 4.86u(k-1) + 3.24u(k-2) - 0.5u(k-1)^2 - 0.06y(k-1)u(k-1) - 0.6u(k-2)^2$$

$y(k)$: DSt (measurement of geomagnetic storm activities)

$u(k)$: VBs (V x Bs)

$f_s = 1/3600$ Hz

A novel analysis of the energy storage process of the magnetosphere system: (i) Observations

VBs
u(k)

$$u(k) = \cos\left(\frac{0.02\pi}{3600} \times 3600 \times k\right) = \cos(0.02\pi k), \quad k = 0, \dots, 6000.$$

$$u(k) = \cos(0.02\pi k) + \cos(0.04\pi k) + \cos(0.06\pi k), \quad k = 0, \dots, 6000$$

$$u(k) = \frac{A}{\pi} \frac{\sin(0.06\pi(k-3000)) - \cos(0.02\pi(k-3000))}{3600(k-3000)}, \quad k = 0, \dots, 6000$$

VBs
frequencies

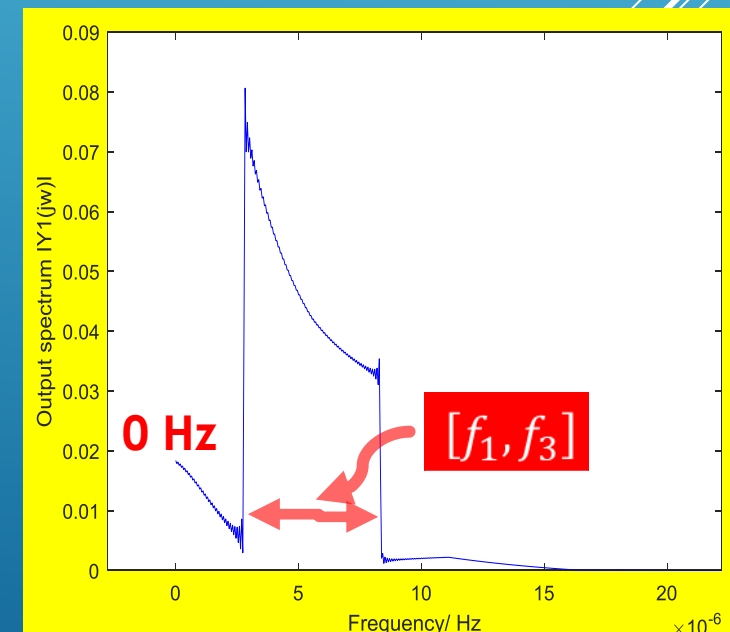
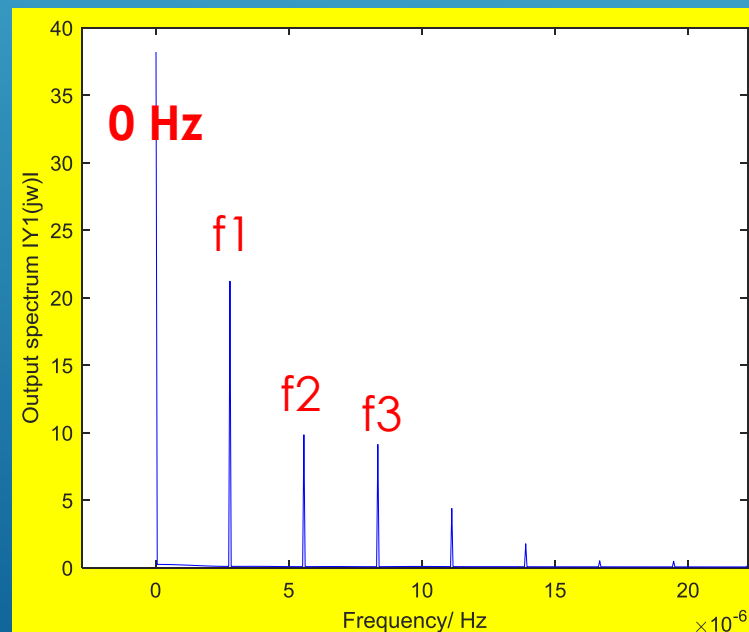
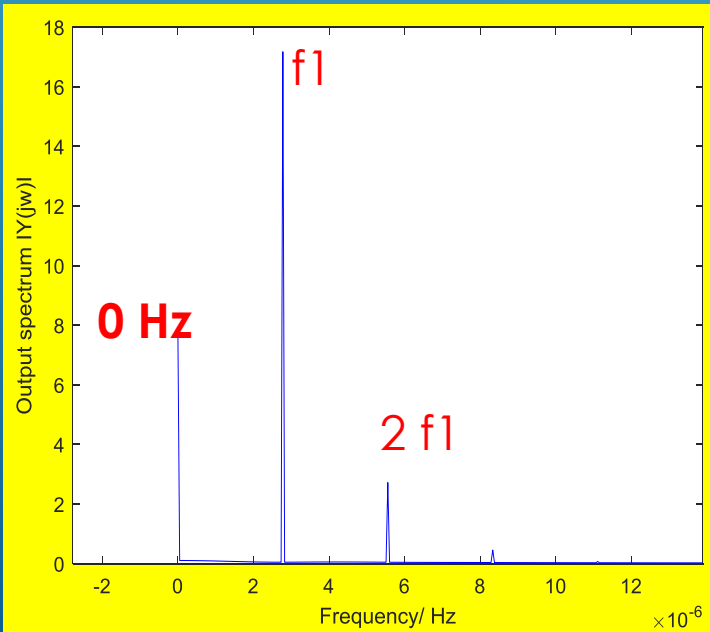
$$f_1 = \frac{0.01}{3600} \text{ Hz}$$

$$f_1 = \frac{0.01}{3600} \text{ Hz}, \quad f_2 = \frac{0.02}{3600} \text{ Hz}, \quad f_3 = \frac{0.03}{3600} \text{ Hz}$$

$$f \in [f_1, f_3] = \left[\frac{0.01}{3600}, \frac{0.03}{3600} \right] \text{ Hz}$$

Dst
Spec
-trum

$Y(j2\pi f)$



A novel analysis of the energy storage process of the magnetosphere system: (ii) Analysis

NARX model of the magnetosphere dynamics

$$y(k) = 0.953y(k-1) - 4.86u(k-1) + 3.24u(k-2) - 0.5u(k-1)^2 - 0.06y(k-1)u(k-1) - 0.6u(k-2)^2$$



Generalized Associated Linear Equations (GALEs) of the NARX model

$$\begin{cases} y_1(k) = 0.953y_1(k-1) - 4.86u(k-1) + 3.24u(k-2) \\ y_2(k) = 0.953y_2(k-1) - 0.5u^2(k-1) \\ \quad - 0.06y_1(k-1)u(k-1) - 0.6u^2(k-2) \\ y_3(k) = 0.953y_3(k-1) - 0.06y_2(k-1)u(k-1) \end{cases}$$



$$y_1(k), y_2(k), y_3(k), \dots$$

components of $y(k)$ contributed by different orders nonlinearity

$$Y_1(j2\pi f)$$

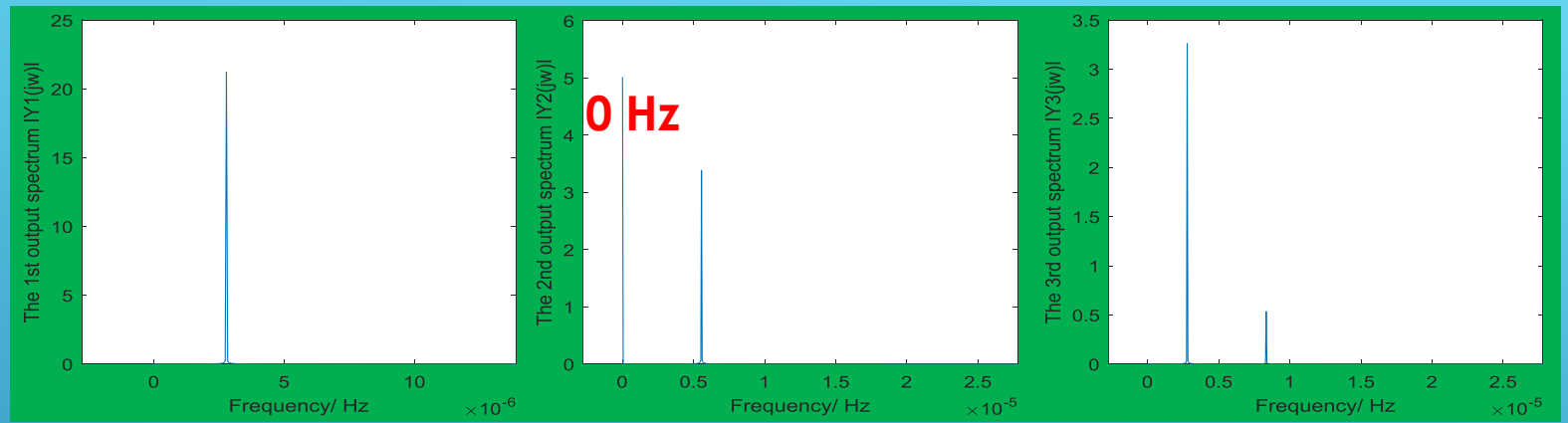
$$Y_2(j2\pi f)$$

$$Y_3(j2\pi f)$$

$$u(k) = \cos\left(\frac{0.02\pi}{3600} \times 3600 \times k\right) = \cos(0.02\pi k), \quad k = \dots, 6000.$$

$$f_1 = \frac{0.01}{3600} \text{ Hz}$$

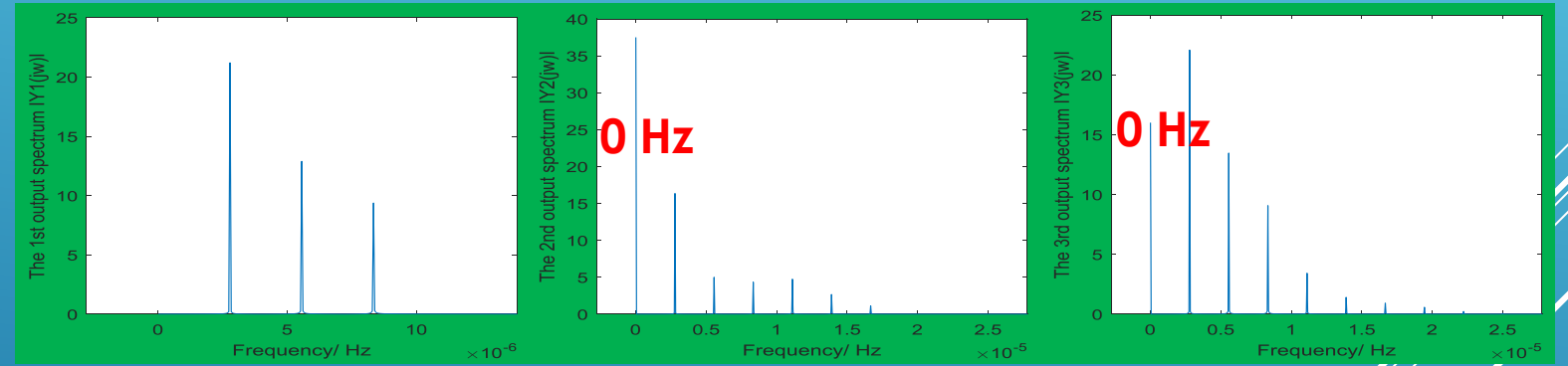
Case 1



$$u(k) = \cos(0.02\pi k) + \cos(0.04\pi k) + \cos(0.06\pi k), \quad k = 0, \dots, 6000$$

$$f_1 = \frac{0.01}{3600} \text{ Hz}, \quad f_2 = \frac{0.02}{3600} \text{ Hz}, \quad f_3 = \frac{0.03}{3600} \text{ Hz}$$

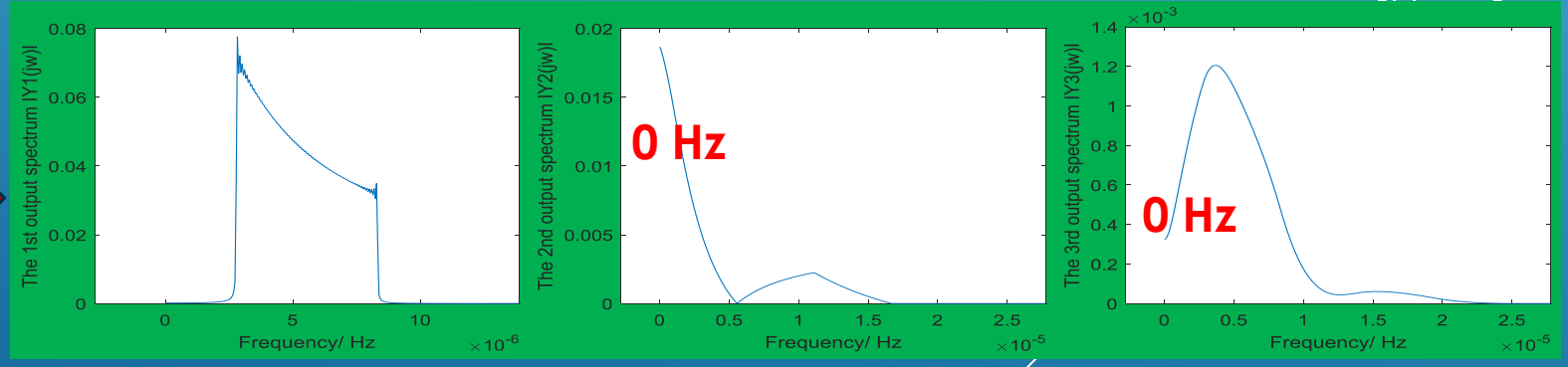
Case 2



$$u(k) = \frac{A \sin(0.06\pi(k-3000)) - \cos(0.02\pi(k-3000))}{\pi \cdot 3600(k-3000)}, \quad k = 0, \dots, 6000$$

$$f \in [f_1, f_3] = \left[\frac{0.01}{3600}, \frac{0.03}{3600} \right] \text{ Hz}$$

Case 3



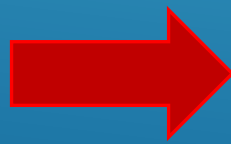
The analysis reveals which order of system nonlinearity contributes to the energy storage

$$\begin{aligned}
 y(k) = & 0.953y(k-1) - 4.86u(k-1) \\
 & + 3.24u(k-2) - \underline{0.5u(k-1)^2} \\
 & - \underline{0.06y(k-1)u(k-1)} \\
 & - \underline{0.6u(k-2)^2}
 \end{aligned}$$



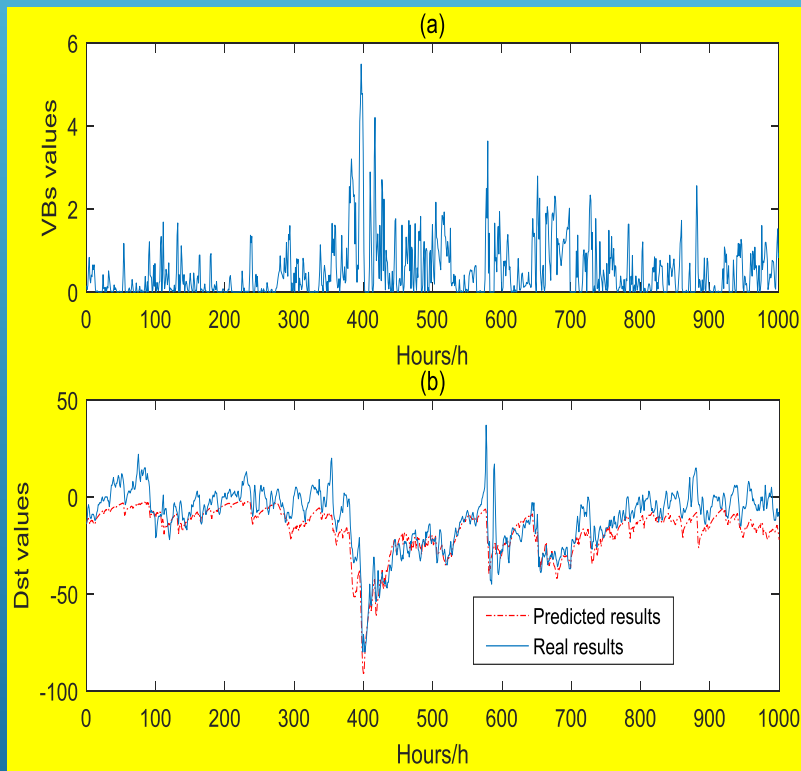
Case	Input frequencies	Order of nonlinearity making contribution to energy storage
Case 1	$f_1 = \frac{0.01}{3600} \text{ Hz}$	2 nd order
Case 2	$f_1 = \frac{0.01}{3600} \text{ Hz}, f_2 = \frac{0.02}{3600} \text{ Hz},$ $f_3 = \frac{0.03}{3600} \text{ Hz}$	2 nd and 3 rd order
Case 3	$f \in [f_1, f_3]$ $= \left[\frac{0.01}{3600}, \frac{0.03}{3600} \right] \text{ Hz}$	2 nd and 3 rd order

$$\begin{cases}
 y_1(k) = 0.953y_1(k-1) - 4.86u(k-1) + 3.24u(k-2) \\
 y_2(k) = 0.953y_2(k-1) - \underline{0.5u^2(k-1)} \\
 \quad - \underline{0.06y_1(k-1)u(k-1)} - \underline{0.6u^2(k-2)} \\
 y_3(k) = 0.953y_3(k-1) - \underline{0.06y_2(k-1)u(k-1)}
 \end{cases}$$

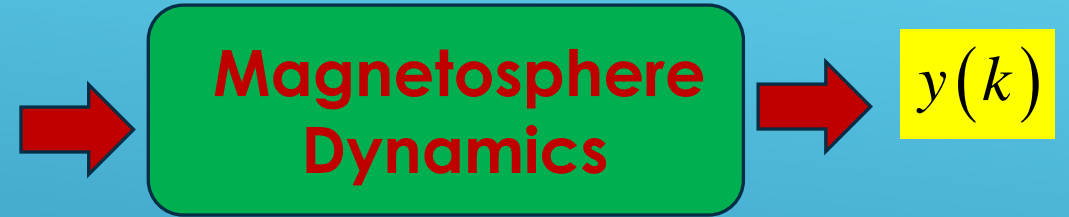


Coefficient 0.06 plays the most significant role making energy storage happen

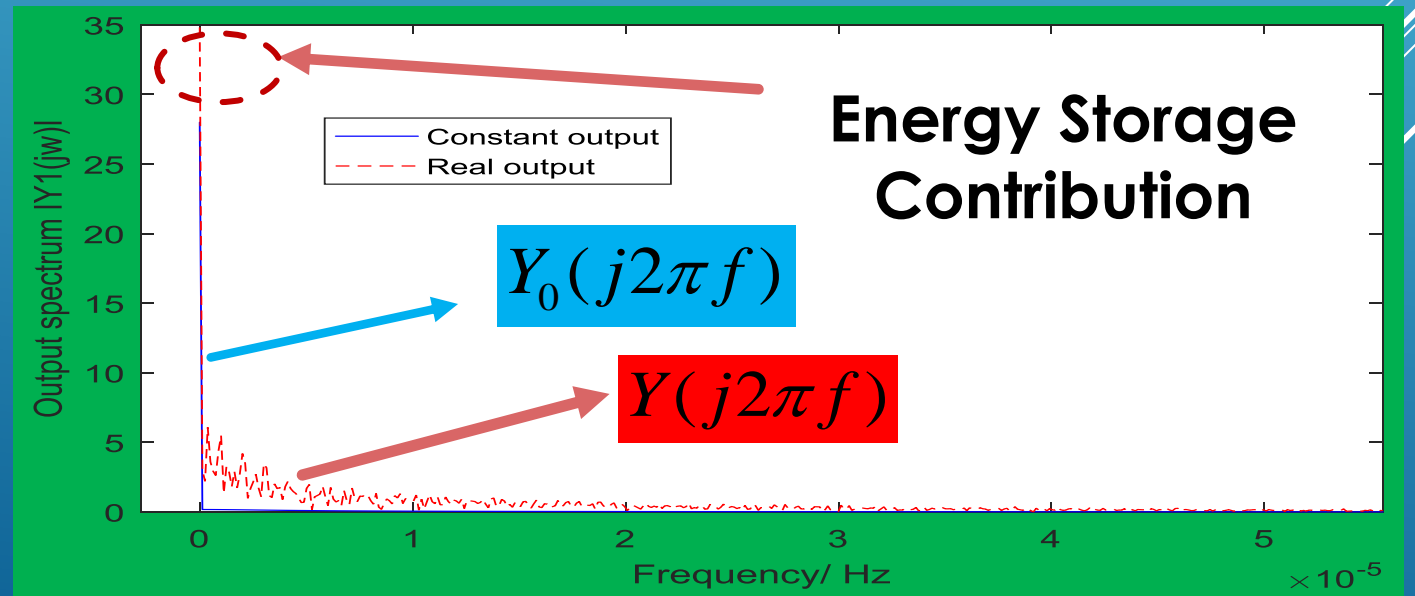
Revelation of energy storage phenomena with the magnetosphere system from satellite data analysis



$$u(k) = C + u_0(k)$$

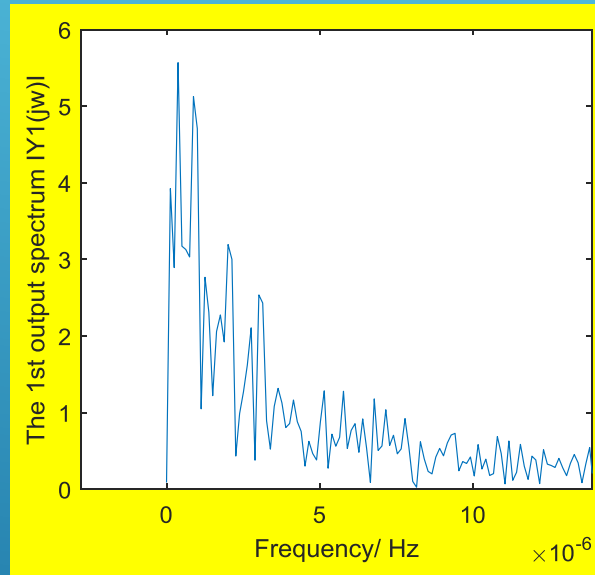


$$u(k) = C$$

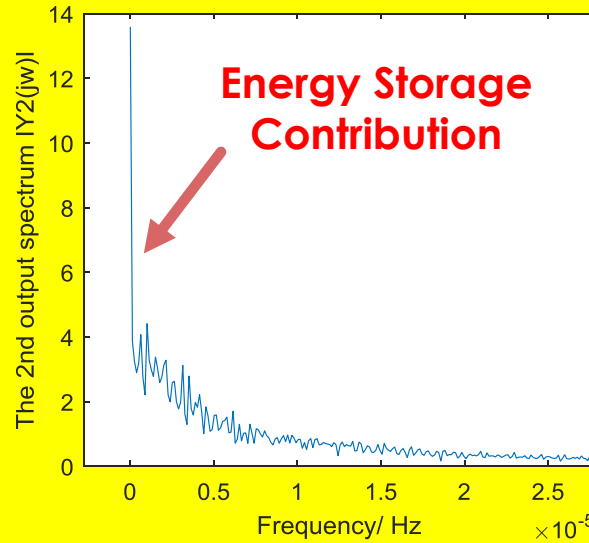


The result of the application of the novel analysis to the satellite data

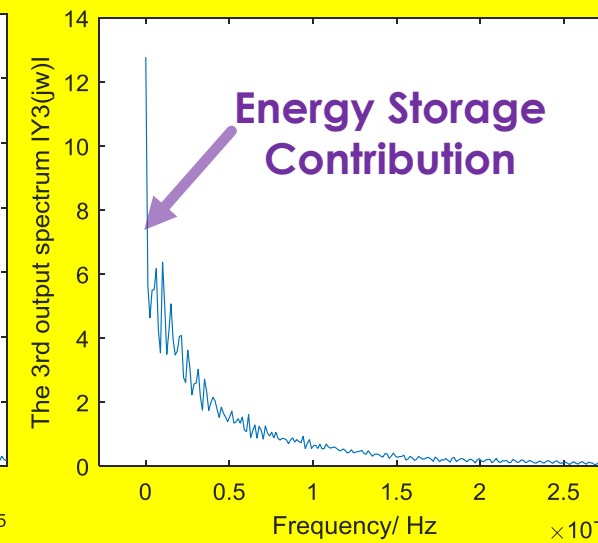
$$Y_1(j2\pi f)$$



$$Y_2(j2\pi f)$$



$$Y_3(j2\pi f)$$



$$\begin{aligned} y(k) = & 0.953y(k-1) - 4.86u(k-1) \\ & + 3.24u(k-2) - \underline{0.5u(k-1)^2} \\ & - \underline{0.06y(k-1)u(k-1)} \\ & - \underline{0.6u(k-2)^2} \end{aligned}$$

Showing again the importance of Coefficient 0.06 in producing the energy storage

An Idea of investigation Of how the energy storage can be affected by relevant physical parameters in space

How the idea could work?

- Build a parametric NARX model of magnetosphere dynamics as

$$\begin{aligned}y(k) = & a_1(P)y(k-1) + b_1(P)u(k-1) \\ & + b_2(P)u(k-2) + c_1(P)u(k-1)^2 \\ & + c_2(P)y(k-1)u(k-1) \\ & + c_3(P)u(k-2)^2\end{aligned}$$

where

$P=[P_1, \dots, P_N]$: Measurable parameters in space that can affect energy storage phenomenon such as. e.g., solar wind density, solar wind pressure, y and x components of the interplanetary magnetic field, and earths seasonal variation etc.

$a_1(P), \dots, c_3(P)$: Known functions of P

- Study how P affects the value of the model coefficients especially $c_2(P)$ hence the energy storage

Conclusions

- **Energy storage** is the phenomena where the **nonlinear magnetosphere dynamics transfer the energy of VBs to Dst at around zero frequency.**
- **A novel nonlinear frequency analysis has been applied to investigate which nonlinear terms in an identified model of magnetosphere dynamics can make most significant contributions to energy storage .**
- **The analyse result could be used to study the effects of other physical parameters in space on energy storage phenomena.**