

Parametric evolution for a deformed cavity

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We consider a classically chaotic system that is described by a Hamiltonian $\mathcal{H}(Q,P;x)$, where (Q,P) describes a particle moving inside a cavity, and x controls a deformation of the boundary. The quantum eigenstates of the system are $|n(x)\rangle$. We describe how the parametric kernel $P(n|m) = |\langle n(x)|m(x_0)\rangle|^2$, also known as the local density of states, evolves as a function of $\delta x = x - x_0$. We illuminate the nonunitary nature of this parametric evolution, the emergence of nonperturbative features, the final nonuniversal saturation, and the limitations of random-wave considerations. The parametric evolution is demonstrated numerically for two distinct representative deformation processes.

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I. INTRODUCTION

A. The local density of states

Consider a system that is described by an Hamiltonian $\mathcal{H}(Q,P;x)$, where (Q,P) are canonical variables and x is a constant parameter. Our interest in this paper is in the case where (Q,P) describe the motion of a particle inside a cavity, and x controls the deformation of the confining boundary. The one-dimensional (1D) version of a cavity, also known as a ‘‘potential well,’’ is illustrated in Fig. 1. However, we are mainly interested in the case of chaotic cavities in $d > 1$ dimensions. Cavities in $d = 2$ dimensions, also known as billiard systems, are prototype examples in the studies of classical and quantum chaos, and we shall use them for the purpose of numerical illustrations.

The eigenstates of the quantized Hamiltonian are $|n(x)\rangle$ and the corresponding eigenenergies are $E_n(x)$. The eigenenergies are assumed to be ordered, and the mean level spacing will be denoted by Δ . We are interested in the parametric kernel

$$P(n|m) = |\langle n(x)|m(x_0)\rangle|^2 = \text{tr}(\rho_n \rho_m). \quad (1)$$

In the equation above $\rho_m(Q,P)$ and $\rho_n(Q,P)$ are the Wigner functions that correspond to the eigenstates $|m(x_0)\rangle$ and $|n(x)\rangle$, respectively. The trace stands for $dQdP/(2\pi\hbar)^d$ integration. The difference $x - x_0$ will be denoted by δx . We assume a dense spectrum, so that our interest is in ‘‘classically small’’ but ‘‘quantum mechanically large’’ energy scales. It is important to realize that the kernel $P(n|m)$ has a well defined classical limit. The classical approximation (see remark [1]) is obtained by using microcanonical distributions instead of Wigner functions.

Fixing n , the vector $P(n|m)$ describes the shape of the n th eigenstate in the $\mathcal{H}_0 = \mathcal{H}(Q,P;x_0)$ representation. By averaging over several eigenstates one obtains the average shape of the eigenstate (ASOE). We can also identify $P(n|m)$ as the local density of states (LDOS), by regarding it as a function of n , where m is considered to be a fixed reference state. In the latter case an average over few m states is assumed. We shall denote the LDOS by $P(r)$, where $r = (n - m)$. The ASOE is just $P(-r)$. Note that the ASOE and the

LDOS are given by the same function. One would have to be more careful with these definitions if \mathcal{H}_0 were integrable while \mathcal{H} nonintegrable.

A few words are in order regarding the definition of the LDOS, and its importance in physical applications. The LDOS, also known as strength function [4–6], describes an energy distribution. Conventionally it is defined as follows:

$$\rho(E) = -\frac{1}{\pi} \langle m | \text{Im} \mathbf{G}(E) | m \rangle = \sum_n P(n|m) \delta(E - E_n), \quad (2)$$

where $\mathbf{G}(E) = 1/(E - \mathcal{H} + i0)$ is the retarded Green function. We are interested in chaotic systems, so it should be clear that our $P(r)$ is related by trivial change of variable ($E \mapsto r$) to the above defined $\rho(E)$. Our $P(r)$ also incorporates an average over the reference state. The LDOS is important in studies of either chaotic or complex conservative quantum systems that are encountered in nuclear physics as well as in atomic and molecular physics. Related applications may be found in mesoscopic physics. Going from \mathcal{H}_0 to \mathcal{H} may signify a physical change of an external field, or switching on of a perturbation, or a sudden change of an effective interaction (as in molecular dynamics [7]). The so-called ‘‘line shape’’ of the LDOS is important for the understanding of the asso-

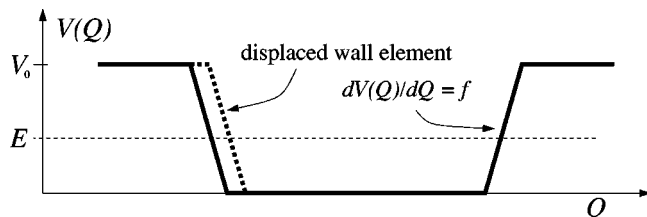


FIG. 1. The shape of a cavity in d dimensions is defined by its $d - 1$ boundary. The confining potential is $V(Q)$. The figure illustrates $V(Q)$ for one dimension well. It also can be regarded as a cross section of the $d > 1$ cavity. The kinetic energy of the particle is $E = \frac{1}{2}mv^2$. The walls of the cavity exert a field of force f on the bouncing particle. The hard wall limit corresponds to $f \rightarrow \infty$ and $V_0 \rightarrow \infty$. For theoretical considerations it is convenient to assume that f and V_0 are large but finite. Mathematically it is also convenient to think of the embedding space as having some huge but finite volume (not illustrated).

ciated dynamics. It is also important to realize that the LDOS is the Fourier transform of the so-called ‘‘survival probability amplitude’’ [7] (see Ref. [8] for concise presentation of this point).

B. Parametric evolution

Textbook [9] formulations of perturbation theory can be applied in order to find the LDOS. Partial summations of diagrams to infinite order can be used in order to get an improved Lorentzian-type approximation. However, most textbooks do not illuminate the limitations and the subtleties which are involved in using the conventional perturbative schemes. It is therefore interesting to take a somewhat different approach to the study of LDOS. The roots of this alternate approach can be traced back to the work of Wigner [10] regarding a simple banded random matrix (BRM) model $\mathcal{H} = \mathbf{E} + \delta x \mathbf{B}$. Here \mathbf{E} is a diagonal matrix whose elements are the ordered energies $\{E_n\}$, and \mathbf{B} is a banded matrix. The study of this model can be motivated by the realization that in generic circumstances it is possible to write $\mathcal{H}(Q, P; x) \approx \mathcal{H}_0(Q, P) + \delta x \mathcal{F}(Q, P)$. Using a simple semiclassical argument [11] it turns out that the matrix representation of any generic \mathcal{F} , in the eigenbasis that is determined by the chaotic Hamiltonian \mathcal{H}_0 , is a banded matrix.

The important ingredient (from our point of view) in the original work by Wigner, is the emphasis on the parametric evolution (PE) of the LDOS. The LDOS describes an energy distribution: For $\delta x = 0$ the kernel $P(r)$ is simply a Kronecker delta function. As δx becomes larger, the width as well as the whole profile of this distribution ‘‘evolves.’’ Wigner has realized that for his BRM model there are three parametric regimes. For very small δx we have the standard perturbative structure where most of the probability is concentrated in $r = 0$. For larger δx we have a Lorentzian lineshape. But this Lorentzian line shape does not persist if we further increase δx . Instead we get a semicircle line shape. Many works about the LDOS have followed [4–6], but the issue of PE has not been further discussed there. The emphasis in those works is mainly on the case where \mathcal{H}_0 is an integrable or noninteracting system, while \mathcal{H} is possibly (but not necessarily) chaotic due to some added perturbation term.

The line of study which is pursued in the present work has been originated and motivated by studies of quantum dissipation [12–14]. Understanding PE can be regarded as a preliminary stage in the analysis of the energy spreading process in driven mesoscopic systems. Note that the LDOS gives the energy redistribution due to a ‘‘sudden’’ (very fast) change of the Hamiltonian. Unlike the common approach for studies of LDOS, we assume both \mathcal{H} and \mathcal{H}_0 to be chaotic. Both correspond to the same parametrically dependent Hamiltonian $\mathcal{H}(Q, P; x)$, and there is nothing special in choosing a particular value $x = x_0$ as a starting point for the PE analysis.

C. Main results

The theory of PE, as discussed in general in Refs. [12–14] and in particular in Refs. [8, 15] takes us beyond the random-matrix-theory considerations of Wigner. There appear five

TABLE I. The parametric scales in the general theory of PE are listed (left column) along with the questions that motivate their introduction. The distribution $P(r)$ may contain perturbative tail regions (for $\delta x \ll \delta x_{\text{prt}}$), and nonperturbative core regions (for $\delta x > \delta x_c^{\text{qm}}$). Nonuniversal (system specific) features may manifest themselves in the core structure for $\delta x \gg \delta x_{\text{NU}}$. In generic examples $\delta x \gg \delta x_{\text{SC}}$ allows a classical approximation for $P(r)$. We are going to explain that only two independent parametric scales survive in the hard wall limit.

δx_c^{cl}	Is it possible to linearize $\mathcal{H}(x + \delta x)$?
δx_c^{qm}	Is it possible to use standard perturbation theory?
δx_{prt}	Do perturbative tail regions survive?
δx_{NU}	Do nonuniversal core features show up?
δx_{SC}	Is it possible to use semiclassical approximation?

(rather than two) different parametric scales (see remark [2]). These are summarized by Table I.

In the present paper we consider cavities with hard walls. We are going to explain that because of the ‘‘hard wall limit’’ there are only *two* independent parametric scales: One is δx_c^{qm} and the others (see remark [3]) coincide with δx_{NU} . Assuming that δx_c^{qm} and δx_{NU} are well separated, it follows that there are *three* distinct parametric regimes in the PE of our system. These are the standard perturbative regime ($\delta x \ll \delta x_c^{\text{qm}}$), the core-tail regime ($\delta x_{\text{NU}} \ll \delta x \ll \delta x_c^{\text{qm}}$), and the nonuniversal regime ($\delta x \gg \delta x_{\text{NU}}$).

The exploration of the three parametric regimes in the PE of a deformed cavity with hard walls is the main issue of the present paper. To the best of our knowledge such detailed exploration has not been practical in the past. We owe our ability to carry out this task to a powerful technique for finding clusters of billiard eigenstates [19, 20]. There are also some secondary issues that we are going to address.

(a) In the strict limit of hard walls the PE becomes non-unitarity. We shall use the 1D well example in order to shed light on this confusing issue. In particular we demonstrate that any truncation of the PE equation leads to false unitarity due to a finite-size edge effect.

(b) For special deformations, namely, those that constitute linear combination of translations, rotations, and dilations, the parametric scales δx_c^{qm} and δx_{NU} coincide. Consequently there is no longer distinct core-tail regime, and the PE features a quite sharp transition from the standard perturbative regime to the nonuniversal regime.

(c) In the nonuniversal regime we demonstrate that our numerical results are in accordance with our theoretical expectation [8]. Namely, the width of the LDOS profile is determined by time-domain semiclassical considerations, rather than by phase-space or random-wave considerations.

(d) The last section puts our specific study in a larger context. We explain why Wigner’s scenario of PE is not followed once hard walls are considered.

II. CAVITY SYSTEM

We consider a particle moving inside a d -dimensional cavity whose volume is V . The kinetic energy of the particle is $E = \frac{1}{2} m v^2$, where m is its mass, and v is its velocity. It is

assumed that this motion is classically chaotic. The ballistic mean free path is l_{bl} . One can use the estimate $l_{\text{bl}} \sim V/A$, where A is the total area of the walls. The associated time scale is $\tau_{\text{bl}} = l_{\text{bl}}/v$.

The penetration distance upon a collision is $l = E/f$, where f is the force that is exerted by the wall. Upon quantization we have an additional length scale, which is the De Broglie wavelength $\lambda_B = 2\pi\hbar/(mv)$. We shall distinguish between the *hard walls* case where we assume $l < \lambda_B \ll l_{\text{bl}}$, and *soft walls* for which $\lambda_B \ll l$. Note that taking $\hbar \rightarrow 0$ implies soft walls.

There is a class of special deformations that are shape preserving. These are generated by translations, rotations, and dilations of the cavity. A general deformation need not preserve the billiard shape nor its volume. We can specify any deformation by a function $D(\mathbf{s})$, where \mathbf{s} specifies the location of a wall element on the boundary (surface) of the cavity, and $D(\mathbf{s})\delta x$ is the normal displacement of this wall element. In many practical cases it is possible to use the convention $|D(\mathbf{s})| \sim 1$. With this convention δx has units of length, and its value has the meaning of typical wall displacement.

The eigenenergies of a particle inside the cavity are in general x dependent, and can be written as $E_n = (\hbar k_n)^2/(2m)$. The mean level spacing is

$$\Delta = \hbar v \times \frac{2\pi}{\Omega_d} \frac{1}{\mathcal{V}} \lambda_B^{d-1} \quad (3)$$

where $\Omega_d = 2\pi, 4\pi, \dots$, for $d = 2, 3, \dots$. In our numerical study we shall consider a quarter stadium with curved edge of radius 1 and straight edge of length 1. The ‘‘volume’’ of the quarter stadium is $\mathcal{V} = 1 + \pi/4$. The ‘‘area’’ of its boundary $A = 4 + \pi/2$ is just the perimeter. We shall look on the parametric evolution of eigenstates around $k \sim 400$ where the mean level spacing in k units is $\tilde{\Delta} = \Delta/(\hbar v) \approx 0.0088$.

III. PARAMETRIC EVOLUTION

Consider the quantum-mechanical state $\psi = |m(x_0)\rangle$. We can write $\psi = \sum_n a_n(x) |n(x)\rangle$. The parametric kernel can be written as $P(n|m) = |a_n(x)|^2$. It is a standard exercise to obtain (from $d\psi/dx = 0$ and differentiating by parts) the following equation for the amplitudes:

$$\frac{da_n}{dx} = -\frac{i}{\hbar} \sum_m \mathbf{W}_{nm}(x) a_m. \quad (4)$$

In order to get $P(n|m)$ one should solve this equation with the initial conditions $a_n(x_0) = \delta_{nm}$. The transitions between levels are induced by the matrix elements

$$\mathbf{W}_{nm} = \frac{i\hbar}{E_n - E_m} \left(\frac{\partial \mathcal{H}}{\partial x} \right)_{nm} \quad (5)$$

and we use the ‘‘gauge’’ convention $\mathbf{W}_{nm} = 0$ for $n = m$. (Only one parameter is being changed and therefore Berry’s phase is not an issue.)

Equation (4) is a possible starting point for constructing a perturbation theory for the PE of $P(r)$ (see Ref. [14] for more details). As an input for this equation we need the matrix elements of $\partial \mathcal{H}/\partial x$. These can be calculated using a simple boundary integral formula [16] whose simplest derivation [14] is as follows: The position of the particle in the vicinity of a wall element is $Q = (z, \mathbf{s})$, where \mathbf{s} is a surface coordinate and z is a perpendicular ‘‘radial’’ coordinate. We take $f = \infty$ so that

$$\frac{\partial \mathcal{H}}{\partial x} = -D(\mathbf{s}) V_0 \delta(z). \quad (6)$$

The logarithmic derivative of the wave function on the boundary is $\varphi(\mathbf{s})/\psi(\mathbf{s})$, where $\varphi(\mathbf{s}) = \mathbf{n} \cdot \nabla \psi$, and \mathbf{n} is a unit vector in the z direction. For $z > 0$ the wave function $\psi(Q)$ is a decaying exponential. If V_0 is large enough, then the exponential decay is fast, and we can treat the boundary as if it were locally flat. It follows that the logarithmic derivative of the wave function on the boundary should be equal to $-\sqrt{2mV_0}/\hbar$. Consequently one obtains the following expression for the matrix elements:

$$\left(\frac{\partial \mathcal{H}}{\partial x} \right)_{nm} = -\frac{\hbar^2}{2m} \oint \varphi_n(\mathbf{s}) \varphi_m(\mathbf{s}) D(\mathbf{s}) ds. \quad (7)$$

In the one-dimensional case the boundary integral is replaced by the sum $\sum_s \varphi_n(s) \varphi_m(s) D(s)$ where $s = 1, 2$ are the two turning points of the potential well.

IV. HARD WALLS AND NONUNITARITY

For the purpose of the following argumentation it is convenient to take $f = \infty$, but to keep V_0 large but finite. Mathematically it is also convenient to think of the embedding space as having some huge but finite volume. We would like to illuminate a subtlety which is associated with the hard wall limit $V_0 \rightarrow \infty$. For any finite V_0 the parametric kernel satisfies

$$\sum_n P(n|m) = p_{\text{total}} \quad (8)$$

with $p_{\text{total}} = 1$. This follows from the fact that $|n(x)\rangle$ is a complete orthonormal basis for any x . However, for hard walls ($V_0 = \infty$) this statement is not true. This implies that for hard walls the PE is nonunitary. We are going to explain this point below.

Let us denote the volume of the original cavity by \mathcal{V}_0 and of the deformed cavity by \mathcal{V} . The volume shared by the deformed and the undeformed cavities will be denoted by $\mathcal{V}_0 \cap \mathcal{V}$ and we shall use the notation $\eta = (\mathcal{V}_0 \cap \mathcal{V})/\mathcal{V}_0$. For the purpose of the following argument let us consider a reference state m whose energy E_m is well below V_0 . Let us also assume that the wall displacement is large compared to De Broglie wavelength. Consequently the expression for $P(n|m)$ has the following semiclassical structure:

$$P(n|m) \approx \eta \times f(E_n - E_m) + (1 - \eta) \times g(E_n - E_c), \quad (9)$$

where $E_c = (\mathcal{V}_0 + E_m) \sim \mathcal{V}_0$. The above result can be deduced by assuming that the wave functions look ergodic in space, but still that they are characterized by a well-defined *local* wavelength. An equivalent derivation is obtained by using the phase-space picture of Ref. [8]. Both f and g in the above expression have unit normalization, and therefore $p_{\text{total}} = 1$ for any finite \mathcal{V}_0 . However, for hard walls ($\mathcal{V}_0 = \infty$) we have $E_c = \infty$ and therefore $p_{\text{total}} = \eta$. We may say the the operation of taking the hard wall limit does not commute with the summation in Eq. (8). An analogous statement can be derived regarding the summation $\sum_m P(n|m)$, with the respective definition $\eta = (V_0 \cap V)/V$.

The correctness of the above observation becomes less trivial if we consider Eq. (4) with expressions (5) and (7) substituted for the matrix elements. Looking at Eq. (5) with Eq. (7) it seems as if the matrix \mathbf{W}_{nm} is Hermitian, and therefore should generate unitary PE. But this statement is mathematically correct only for (any) finite truncation N of the PE equation. For $N = \infty$ the matrix \mathbf{W}_{nm} becomes non-Hermitian. It turns out that for any finite N , there is a pile-up of probability in the edges of the spreading profile, due to finite-size effect. We shall demonstrate this effect in the next section using a simple 1D example. In other words, if we solve Eq. (4) for hard-walled cavity, we get as a result Eq. (9) with $E_c = E_N$. For $N = \infty$ we get $E_c = \infty$ and therefore $p_{\text{total}} = \eta$ in accordance with the conclusion of the previous paragraph.

Thus if either $V_0 < \infty$ or $N < \infty$ then we have unitary PE. But for hard walls, meaning $V_0 = \infty$ with $N = \infty$, we have nonunitary PE. The lost probability is associated with the second term in Eq. (9). This term is peaked around a high energy E_c . For hard walls $E_c = \infty$ and consequently some probability is lost. The above picture is supported by the simple pedagogical example of the next section.

V. PARAMETRIC EVOLUTION FOR A 1D BOX

Consider a 1D box with hard walls, where the free motion of the particle is within $0 < Q < \mathbf{a}$. The eigenstates of the Hamiltonian are

$$|n(\mathbf{a})\rangle \rightarrow (-1)^n \sqrt{\frac{2}{\mathbf{a}}} \sin(k_n Q), \quad (10)$$

where $k_n = n \times (\pi/\mathbf{a})$ is the wave number, and $n = 1, 2, \dots$, is the level index. The phase factor $(-1)^n$ has been introduced for convenience. We consider now the parametric evolution as a function of \mathbf{a} . One easily obtains

$$\langle n(\mathbf{a}) | m(\mathbf{a}_0) \rangle = (-1)^n \sqrt{\eta} \frac{\sin(\pi \eta n)}{\pi} \frac{2m}{\eta^2 n^2 - m^2}, \quad (11)$$

where $\eta \equiv \mathbf{a}_0/\mathbf{a}$ is assumed to be smaller than 1, corresponding to expansion of the box. The probability kernel is $P(n|m) = |\langle n|m \rangle|^2$. One can verify that the parametric evolution in the $\mathbf{a}_0 \mapsto \mathbf{a}$ direction is unitary, meaning that $\sum_n P(n|m) = 1$. On the other hand, in the $\mathbf{a} \mapsto \mathbf{a}_0$ direction the

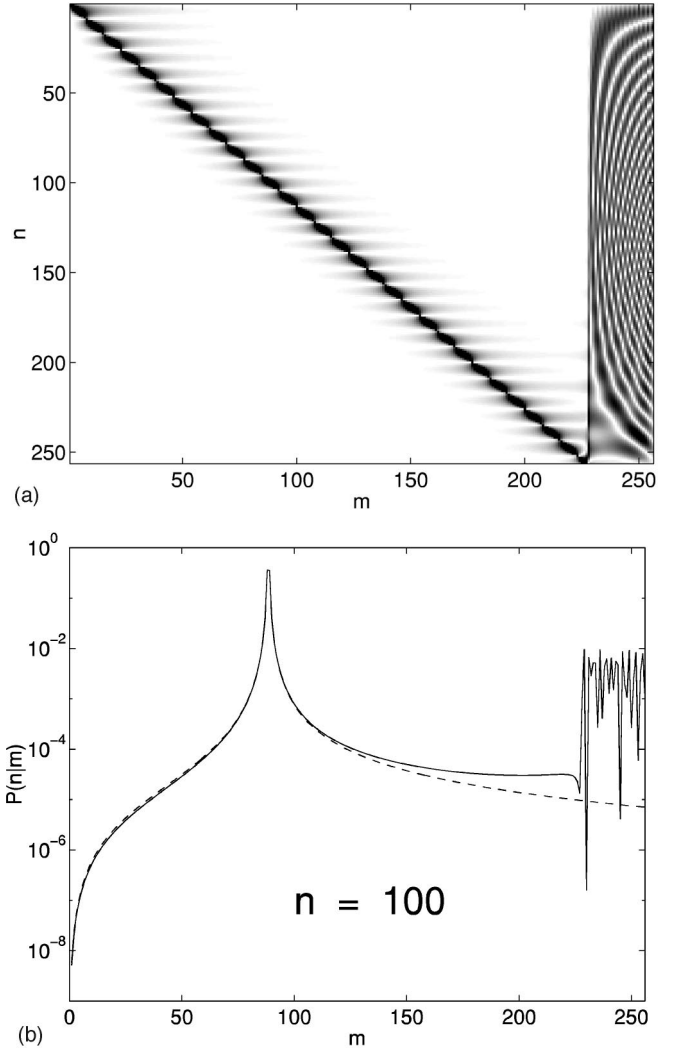


FIG. 2. (a) An image of the kernel $P(n|m)$ for 13% expansion of the box (i.e., $\mathbf{a}/\mathbf{a}_0 = 1.13$). The kernel has been calculated numerically using Eq. (13) with finite truncation $N = 256$. (b) The profile of a representative row of $P(n|m)$. The dashed line is the $N = \infty$ analytical result using Eq. (11).

parametric evolution is nonunitary, because $\sum_m P(n|m) = \eta$. The profile of $P(n|m)$ for fixed n is illustrated by a dashed line in Fig. 2.

We can restore unitarity by making V_0 large but finite. In such case, a variation of the above calculation leads to the following picture: Consider the overlap of a reference level $n(\mathbf{a})$ with the levels $m(\mathbf{a}_0)$. As in the case $V_0 = \infty$ there is a probability η which is located in the levels whose energies are $E_m \approx E_n$. But now the “lost” probability $(1 - \eta)$ is located in the levels whose energies are $E_m \approx E_n + V_0$. Thus we have $\sum_m P(n|m) = 1$ rather than $\sum_m P(n|m) = \eta$.

We again consider the case $V_0 = \infty$. The normal derivative on the boundary is $\varphi_n(\mathbf{a}) = \sqrt{(2/\mathbf{a})} k_n$. Hence we can easily get the following result:

$$\frac{1}{\hbar} \mathbf{W}_{nm} = \frac{-i}{k_n^2 - k_m^2} \varphi_n(\mathbf{a}) \varphi_m(\mathbf{a}) = -i \frac{1}{\mathbf{a}} \frac{2nm}{n^2 - m^2}. \quad (12)$$

It is more convenient to use $\alpha = \ln(\mathbf{a})$ for parametrization. Hence the equation that describes the parametric evolution is

$$\frac{da_n}{d\alpha} = - \sum_m \frac{2nm}{n^2 - m^2} a_m. \quad (13)$$

For any finite truncation $N < \infty$ this equation manifestly generates unitary parametric evolution. It is only for $N = \infty$ that it becomes equivalent to the nonunitary evolution of the 1D box. Again, one can wonder where the ‘‘lost’’ probability is located if $N < \infty$. The answer is illustrated in Fig. 2. We see that the ‘‘lost’’ probability piles up at the edge of the (truncated) tail.

VI. MATRIX ELEMENTS FOR CHAOTIC CAVITY

It is possible to use semiclassical considerations [11] in order to determine the band profile of the matrix Eq. (7). The application to the cavity example has been introduced in Ref. [14], and numerically demonstrated in Ref. [17]. The accuracy of this semiclassical estimate is remarkable. Here we summarize the recipe. First one should generate a very long (ergodic) trajectory, and define for it the fluctuating quantity

$$\mathcal{F}(t) = - \frac{\partial \mathcal{H}}{\partial x} = \sum_{\text{col}} 2mv \cos(\theta_{\text{col}}) D_{\text{col}} \delta(t - t_{\text{col}}), \quad (14)$$

where t_{col} is the time of a collision, D_{col} stands for $D(\mathbf{s})$ at the point of the collision, and $v \cos(\theta_{\text{col}})$ is the normal component of the particle’s velocity. Each delta spike (for soft walls it is actually a narrow rectangular spike) corresponds to one collision. Now one can calculate the correlation function $C(\tau)$ of the fluctuating quantity $\mathcal{F}(t)$, and its Fourier transform $\tilde{C}(\omega)$. The semiclassical estimate for the band profile is

$$\left\langle \left| \left(\frac{\partial \mathcal{H}}{\partial x} \right)_{nm} \right|^2 \right\rangle \approx \frac{\Delta}{2\pi\hbar} \tilde{C} \left(\frac{E_n - E_m}{\hbar} \right) \quad (15)$$

Reference [18] contains a systematic study of the function $\tilde{C}(\omega)$. For large ω , meaning $\omega \gg 1/\tau_{\text{bl}}$, one can use

$$\tilde{C}(\omega) \approx 2m^2 v^3 \langle |\cos \theta|^3 \rangle \frac{1}{V} \oint [D(\mathbf{s})]^2 ds, \quad (16)$$

where the geometric factor is $\langle |\cos(\theta)|^3 \rangle = 1, 4/3\pi, \dots$, for $d = 1, 2, \dots$. A lengthy calculation [14] reveals that Eq. (15) with (16) substituted, is an exact global result if we could assume that the cavity eigenfunctions look similar to ‘‘random waves,’’ and that different wave functions are uncorrelated. However, it turns out that to take this random wave result as a global approximation is an oversimplification. For $\omega \ll 1/\tau_{\text{bl}}$, using the semiclassical recipe and assuming strongly chaotic cavity, one obtains

$$\tilde{C}(\omega) \approx \tilde{C}(\infty) \times (\tau_{\text{bl}} \omega)^\gamma \quad (17)$$

with $\gamma = 4$ for dilations and translations, $\gamma = 2$ for rotations, and $\gamma = 0$ for normal deformations. We use the term ‘‘special deformations’’ [18] in order to distinguish those deforma-

tions that have the property $\tilde{C}(\omega) \rightarrow 0$ in the limit $\omega \rightarrow 0$. Any combination of dilations, translations, and rotations is a special deformation. Around the bouncing frequency ($\omega \sim 1/\tau_{\text{bl}}$) the function $\tilde{C}(\omega)$ typically displays some nonuniversal (system and deformation specific) structure. This is true for any typical deformation, but for some deformations the nonuniversal features are more pronounced. If the cavity has bouncing ball modes, we may get also a modified (nonuniversal) behavior in the small frequency limit. For the purpose of general discussion it is convenient to assume that the interpolation between Eqs. (17) and (16) is smooth, but in actual numerical calculation the actual $\tilde{C}(\omega)$ is computed (see below).

As a numerical example we have picked the stadium billiard. We have found the eigenstates of a desymmetrized (quarter) stadium as described in Ref. [17]. We have selected those eigenstates whose eigenenergies k_n are in the vicinity of $k = 400$. Our two representative deformations are (a) rotation around the stadium center and (b) generic (nonspecial) deformation involving the curved edge. In the latter case the curved edge of the quarter stadium ($0 < s < \pi/2$) is pushed outwards with $D(s) = [\cos(s)]^2$, while for the straight edges $D(s) = 0$. (The corner $s = 0$ is the 90° intersection of the curved edge with the long straight edge). The respective band profiles are displayed in Fig. 3. The band profile has been defined as

$$B(\kappa) = \frac{1}{4k^2} \left\langle \left| \oint \varphi_n(\mathbf{s}) \varphi_m(\mathbf{s}) D(\mathbf{s}) ds \right|^2 \right\rangle, \quad (18)$$

where $\kappa = (k_n - k_m)$ is the distance from the diagonal. Note that $B(\kappa)$ is just a scaled version of the semiclassical $\tilde{C}(\omega)$ as implied by Eq. (15) with (7). The remarkable agreement of $B(\kappa)$ with the semiclassical calculation has been demonstrated in Refs. [17,18].

It is important to realize that in the hard wall limit (which is assumed here) the matrix $(\partial \mathcal{H} / \partial x)_{nm}$ is not a banded matrix. It would become banded if we were assuming *soft* walls. For soft walls $\tilde{C}(\omega)$ becomes vanishingly small for $\omega \gg v/l$. The bandwidth in energy units is $\Delta_b = \hbar v/l$, and in dimensionless units it is

$$b = \frac{\Delta_b}{\Delta} = \frac{\mathcal{V}}{l \lambda_B^{d-1}}. \quad (19)$$

Unless stated otherwise we have $b = \infty$.

VII. PARAMETRIC EVOLUTION—NUMERICAL RESULTS

The parametric evolution of $P(r)$ for rotation and for generic deformation of the stadium is illustrated by the images of Fig. 4 and by the plots of Fig. 5 and Fig. 6. The calculation of each $P(r)$ profile is carried out as follows: Given δx we use the method which is described in Ref. [20] in order to calculate the matrix $P(n|m)$. Then we plot the elements of $P(n|m)$ versus $\kappa = [k_n(x) - k_m(0)]$. In order to obtain the

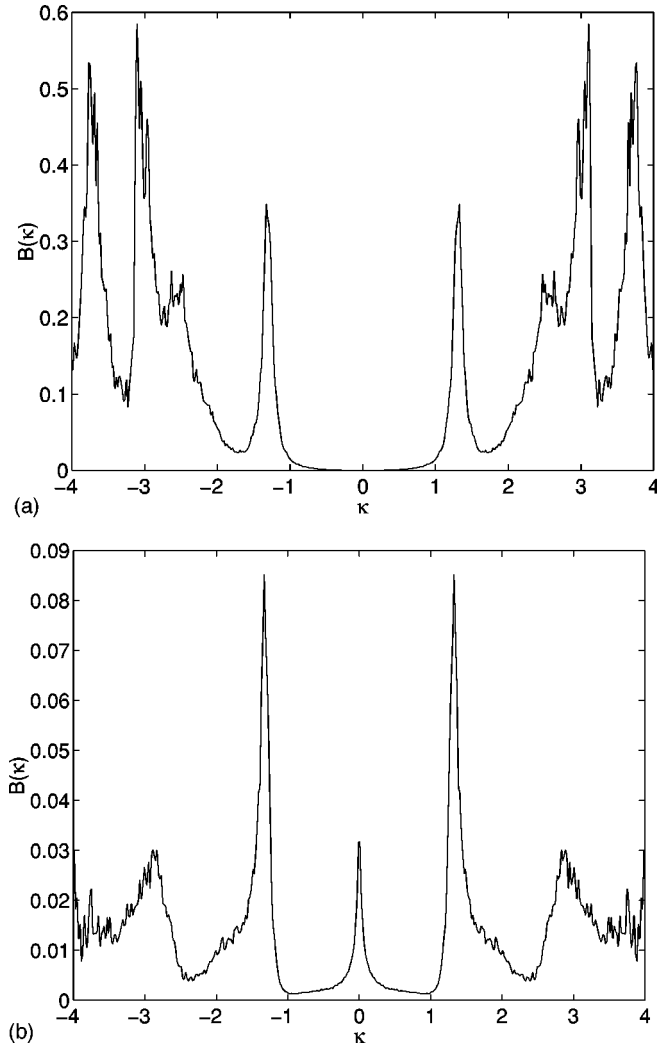


FIG. 3. Band profiles for deformations of the quarter stadium as defined by Eq. (18). (a) Rotation around the stadium center. (b) Generic (nonspecial) deformation involving displacement of the curved edge. It is important to notice that for the special deformation we have $\tilde{C}(\omega) \rightarrow 0$ in the limit $\omega \rightarrow 0$.

average profile the plot is smeared using standard procedure (see remark [21]). The transformation from κ to $r = (n - m)$ is done using the relation (see remark [21]):

$$\kappa = \tilde{\Delta} \cdot r - \frac{1}{d} k \times \frac{\delta \mathcal{V}}{\mathcal{V}}. \quad (20)$$

Above $\tilde{\Delta}$ is the mean level spacing of the $\{k_n\}$ spectrum, and $\delta \mathcal{V}$ is the volume change that is associated with the deformation (it is approximately proportional to δx). If the deformation is volume preserving (as in the case of rotation) then the second term equals zero. But for the generic deformation that we have picked in our second numerical example, the volume is not preserved, and the systematic ‘‘downward’’ shift of the levels should be taken into account.

Looking first at the case of rotation, we clearly see two parametric regimes: The standard perturbative regime ($\delta x < 0.2$) and the nonuniversal regime ($\delta x > 0.2$). Let us clarify

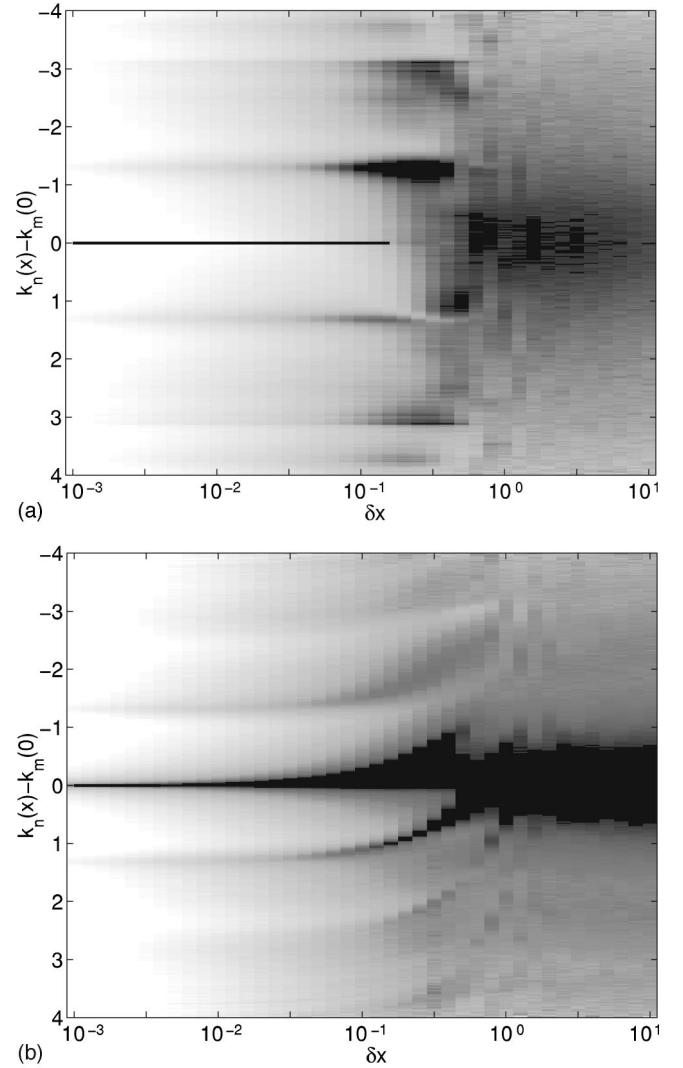


FIG. 4. Each column is an image of $P(r)$ versus κ for a different value of δx . There are 41 columns. The value $\delta x \sim 1$ corresponds roughly to λ_B displacement of the boundary. (a) is for rotation and (b) is for generic deformation. Note that the $r=0$ component is excluded from the image. Instead we have plotted over the image an $r=0$ line wherever this component contains more than 50% of the probability. In the lower figure this line cannot be resolved from the developing core region.

this observation. We see that for $\delta x < 0.2$ most of the probability is well concentrated in $r=0$. This implies that we can use standard perturbation theory in order to estimate the $r=0$ probabilities. On the other hand, for $\delta x > 0.2$ the perturbative nature of $P(r)$ is destroyed. Now $P(r)$ becomes smoother, and eventually (for $\delta x > 0.5$) there is a very good fitting with Lorentzian (see lower plot in Fig. 5).

The qualitative explanation for the Lorentzian profile is as follows. For $\delta x > 0.5$ the typical displacement of the walls is of the order of λ_B . Therefore the $|n(x)\rangle$ eigenstates become uncorrelated with the $|m(0)\rangle$ eigenstates. Consequently $P(r)$ becomes δx independent. The Lorentzian profile agrees with the assumption of uncorrelated random waves as explained in the Appendix.

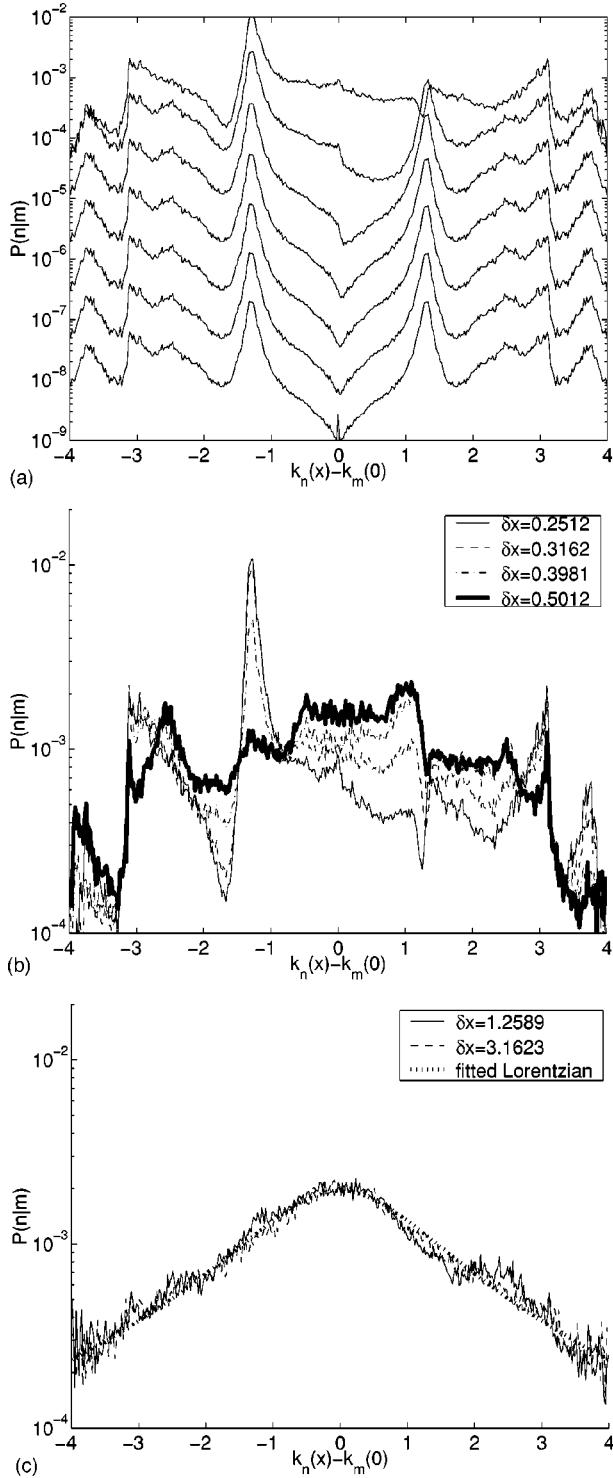


FIG. 5. Representative plots of $P(r)$ for the case of rotation. (a) is for $0.0010 \leq \delta x \leq 0.2512$. The $r=0$ component is excluded. In (b) the fitted Lorentzian is indistinguishable from the actual profile.

Let us now look at the case of generic deformation. Here we clearly see three parametric regimes: The standard perturbative regime ($\delta x < 0.004$), the core-tail regime ($0.004 < \delta x < 0.2$), and the nonuniversal regime ($\delta x > 0.2$). Let us clarify this observation. As in the case of rotation there is a standard perturbative regime ($\delta x < 0.004$) where most of the

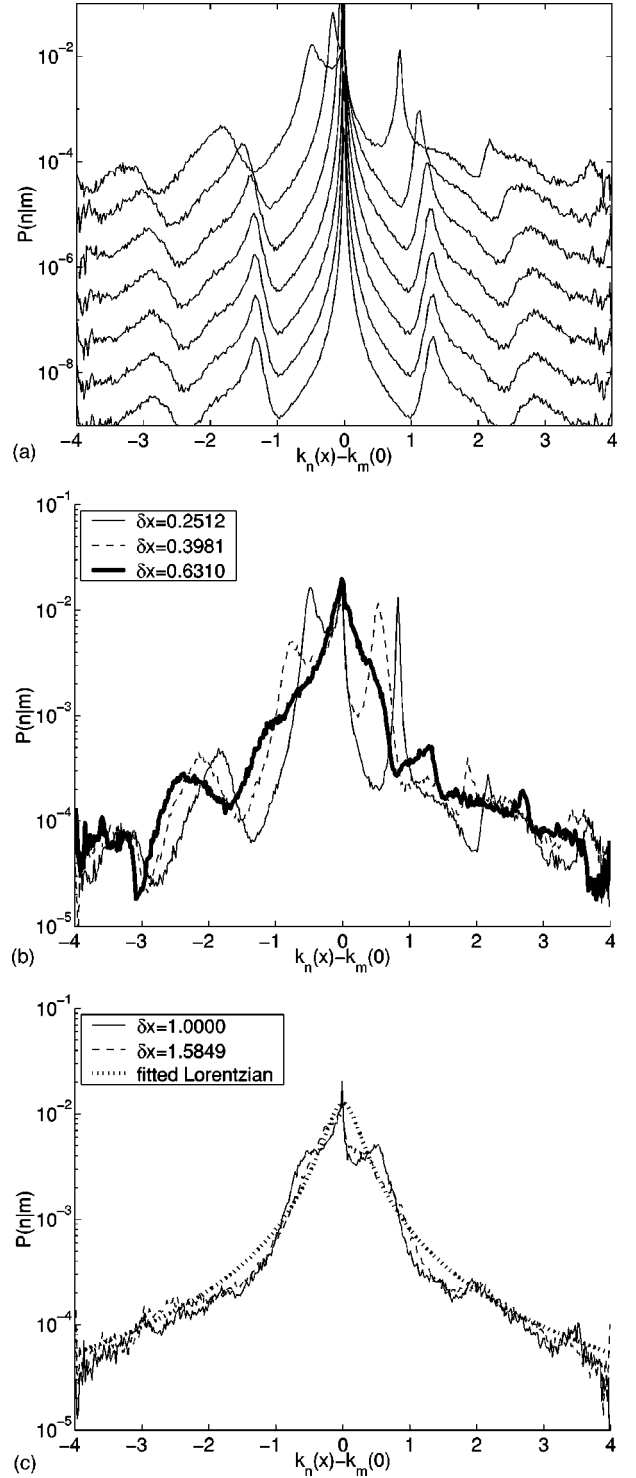


FIG. 6. Representative plots of $P(r)$ for the case of generic deformation. (a) is for $0.0010 \leq \delta x \leq 0.2512$. In (b) a fitted Lorentzian is overlaid for the purpose of comparison.

probability is well concentrated in $r=0$. For larger deformation, namely for $\delta x > 0.004$, standard perturbation theory is no longer applicable because the $n=m$ level is mixed nonperturbatively with other (neighboring) levels. As a result $P(r)$ contains a nonperturbative ‘‘core’’ component. However, for $0.004 < \delta x < 0.2$ we definitely do not get a Lorentzian

zian. Rather the tails of $P(r)$ keep growing in the same way as in the standard perturbative regime.

In the case of the generic deformation, as in the case of rotation, we enter the nonuniversal regime, and eventually (for $\delta x > 0.6$) we get a smooth Lorentzian-like distribution. However, the Lorentzian-like distribution is not identical to that of the rotation case. Also the similarity to proper Lorentzian is far from being satisfactory (see lower plot in Fig. 6). This means that the random-wave picture of the Appendix is an oversimplification.

In the following sections we are going to summarize the theoretical considerations [8] that explain the observed parametric scenario. In particular we are going to illuminate the way in which nonperturbative features emerge; to clarify the crossover to the nonuniversal regime; and to explain the specific nature of the nonuniversal distribution.

VIII. THE STANDARD PERTURBATIVE REGIME

Standard perturbation theory gives the following first order expression for the LDOS

$$P(n|m) \approx \delta_{nm} + \delta x^2 \frac{|\partial \mathcal{H} / \partial x_{nm}|^2}{(E_n - E_m)^2}. \quad (21)$$

This expression is most straightforwardly obtained by inspecting Eqs. (4) and (5). We can define the (total) transition probability as

$$p(\delta x) = \sum_{r \neq 0} P(r). \quad (22)$$

Using Eq. (21) combined with (15) we get the following estimate:

$$p(\delta x) \approx \delta x^2 \times \frac{1}{\hbar^2} \int_{|\omega| > \Delta/\hbar} \frac{d\omega}{2\pi} \frac{\tilde{C}(\omega)}{\omega^2}. \quad (23)$$

Standard perturbation theory is applicable as long as $p(\delta x) \ll 1$. This can be converted into an equivalent inequality $\delta x \ll \delta x_c^{\text{qm}}$. By this definition δx_c^{qm} is the parametric deformation which is needed in order to mix the initial level m with other levels $n \neq m$.

If we use Eq. (23) for a special deformation, then we have $\gamma > 1$, and consequently the integral is not sensitive to the exclusion of the $|r| < \Delta$ region. As a result we have $\delta x_c^{\text{qm}} \propto \hbar$. Using Eq. (17) with (16) we get

$$\delta x_c^{\text{qm}}|_{\text{special}} = \left(l_{\text{bl}} \frac{1}{\gamma} \oint [D(\mathbf{s})]^2 d\mathbf{s} \right)^{-1/2} \times \lambda_B. \quad (24)$$

In the case of generic deformation, Eq. (24) is not valid because the value of the integral in Eq. (23) is predominantly determined by the $\omega \sim \Delta/\hbar$ lower cutoff rather than by an effective $\omega \sim 1/\tau_{\text{bl}}$ lower cutoff. As a result one obtains

$$\delta x_c^{\text{qm}}|_{\text{generic}} = \left(l_H \frac{1}{\gamma} \oint [D(\mathbf{s})]^2 d\mathbf{s} \right)^{-1/2} \times \lambda_B, \quad (25)$$

where $l_H = v t_H$ is the length which is associated with the Heisenberg time $t_H = 2\pi\hbar/\Delta$.

It is illuminating to use the convention $D(\mathbf{s}) \sim 1$, such that δx measures the typical displacement of a wall element. With this convention we get from Eqs. (24),(25) the following:

$$\delta x_c^{\text{qm}}|_{\text{special}} \approx \lambda_B, \quad (26)$$

$$\delta x_c^{\text{qm}}|_{\text{generic}} = \left(\frac{\lambda_B^{d-1}}{A_{\text{eff}}} \right)^{1/2} \times \lambda_B, \quad (27)$$

$$\delta x_c^{\text{qm}}|_{\text{diffractive}} \sim \lambda_B. \quad (28)$$

In the generic case the effective area of the deformed boundary A_{eff} may be smaller than the total area A of the boundary. The effective area A_{eff} can be formally defined by comparing Eq. (27) with Eq. (25). Equation (28) has been added for sake of completeness of our presentation. It correspond to the diffractive limit $A_{\text{eff}} \rightarrow \lambda_B^{d-1}$. Note that this limit is beyond the scope of the present study. Thus in the generic case, the wall displacement needed to mix levels is much smaller than λ_B . In the generic case $\delta x_c^{\text{qm}} \propto \hbar^{(d+1)/2}$ rather than $\delta x_c^{\text{qm}} \propto \hbar$. What happens with perturbation theory beyond δx_c^{qm} ? This is the subject of the next section.

IX. THE CORE-TAIL REGIME

For $\delta x > \delta x_c^{\text{qm}}$ standard perturbation theory diverges due to the nonperturbative mixing of neighboring levels on small scale: Once δx becomes of the order of δx_c^{qm} several levels are mixed, and as δx becomes larger, more levels are being mixed nonperturbatively. Consequently it is natural to distinguish between *core* and *tail* regions [13–15]. Most of the spreading probability is contained within the core region, which implies a natural extension of first-order perturbation theory (FOPT): The first step is to transform Eq. (4) into a new basis where transitions within the core are eliminated. The second step is to use FOPT (in the new basis) in order to analyze the core-to-tail transitions. Details of this procedure, which is in the spirit of degenerate perturbation theory, were discussed in Ref. [14]. The most important (and nontrivial) consequence of this procedure is the observation that mixing on small scales does not affect the transitions on large scales. Therefore we have in the tail region $P(n|m) \propto \delta x^2$ rather than $P(n|m) \propto \delta x$. The validity of this observation has been numerically illustrated in Ref. [15].

Following the above reasoning we define the *tail* region as consists of those levels whose ‘‘occupation’’ can be calculated using perturbation theory, while the *core* is the non-perturbative component in the vicinity of $r=0$. Assuming that only one scale characterize the core width, one arrives at the following practical approximation

$$P_{\text{prt}}(r) = \frac{\Delta}{2\pi\hbar} \tilde{C} \left(\frac{E_n - E_m}{\hbar} \right) \frac{\delta x^2}{[\Gamma(\delta x)]^2 + (E_n - E_m)^2}. \quad (29)$$

It is implicit in this definition that $(E_n - E_m)$ should be regarded as a function of $r = (n - m)$. The parameter $\Gamma(\delta x)$ is determined (for a given δx) such that $P_{\text{prt}}(r)$ has a unit normalization. One may say that $\Gamma(\delta x)$ regularizes the behavior around $r = 0$. For generic deformation we get

$$\Gamma(\delta x) \equiv b_0 \Delta \approx \left(\frac{\delta x}{\delta x_c^{\text{qm}}} \right)^2 \times \Delta. \quad (30)$$

The core is defined as the region $|r| < b_0$, and the outer ($|r| > b_0$) regions are the tails. For $\delta x \ll \delta x_c^{\text{qm}}$ we get $b_0 \ll 1$ and the core-tail structure Eq. (29) becomes equivalent to the standard perturbative result Eq. (21). It should be clear that the core-tail structure is a generalization of Wigner's Lorentzian. It is indeed a Lorentzian in the special case of a "flat" band profile.

We turn now to analyze our numerical results. We have verified (see, e.g., Fig. 7) that for $\delta x < 0.2$ we have good agreement with perturbation theory irrespective of whether we have a core component (which is the case for the generic deformation) or not (which is the case for the rotation). As we come closer to $\delta x \sim 0.2$ the agreement becomes worse, and for $\delta x > 0.3$ we have a total collapse of perturbation theory. In Fig. 8 we display the total transition probability $p(\delta x)$ as a function of δx . This plot should be used in order to numerically determine the value of δx_c^{qm} , say as the value where $p(\delta x) = 1/2$.

Figure 8 also displays comparison with the corresponding perturbative calculation [using Eq. (22) with (29)]. The agreement (for $\delta x < \delta x_c^{\text{qm}}$) in case of the rotation is remarkable. For larger values of δx one observes a linear drop of the total probability due to the nonunitary nature of the evolution. (This drop is clearly linear in a linear-linear scale which is not displayed.) The eventual rise of the total transition probability (beyond 1) in case of the generic deformation, reflects numerical errors in the determination of the small κ overlaps. See Ref. [20] for further details.

The lower subfigure in Fig. 8 displays the calculated core width b_0 for the generic deformation. Recall that b_0 is determined, given δx , such that $P_{\text{prt}}(r)$ of Eq. (29) is normalized. Also calculated is the width b_c of the region that contains 50% of the probability. The width b_c is calculated for both $P_{\text{prt}}(r)$ and $P(r)$. Note that $b_c = 1$ in the standard perturbative regime. The determination of δx_c^{qm} for the generic deformation becomes more convenient by using this plot. Also the crossover (at $\delta x_{\text{NU}} \sim 0.2$) to the nonuniversal regime is most pronounced.

X. NONUNIVERSAL REGIME

The validity of Eq. (29) as a global approximation relies on the assumption that the core is characterized only by one scale (which is b_0). But this assumption ceases to be true if δx is large enough. We have explained in Ref. [8] that the width of the *core* defines a "window" through which we can view the "landscape" of the semiclassical analysis. As δx becomes larger, this "window" becomes wider, and eventually some of semiclassical structure is exposed. This is

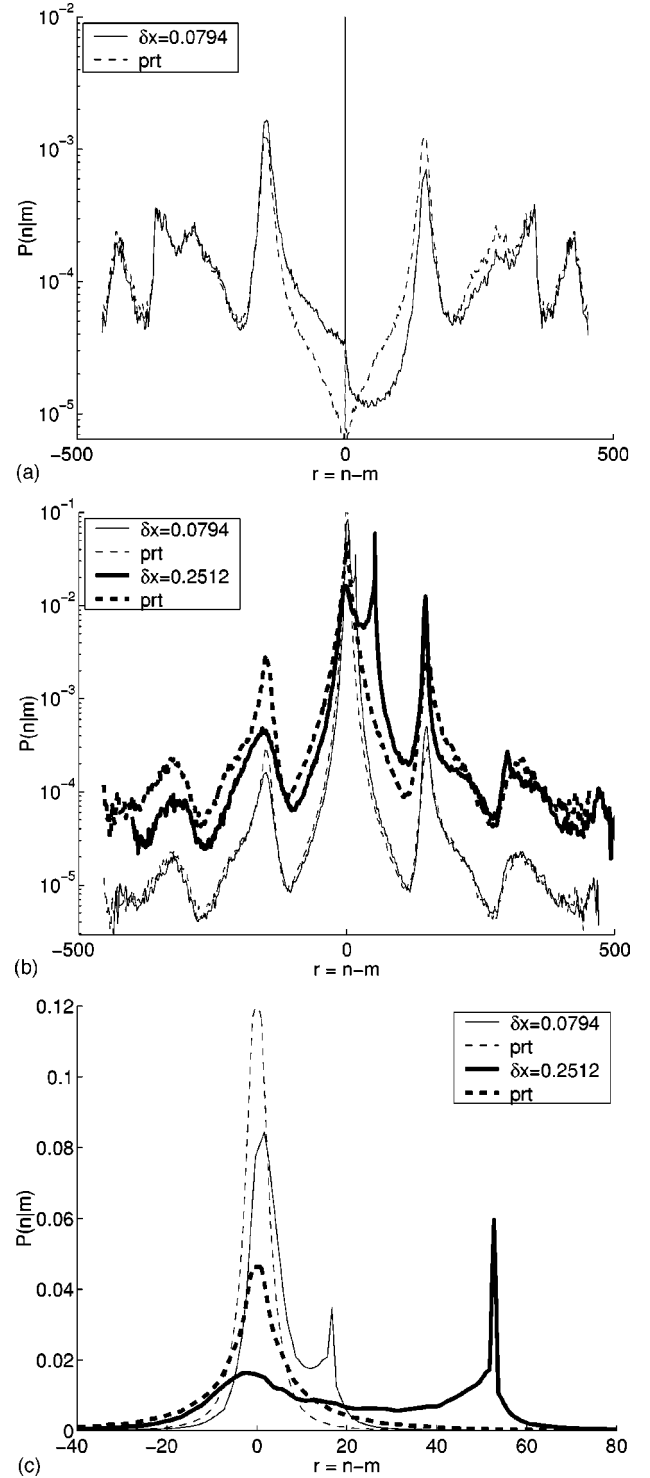


FIG. 7. Comparison between the perturbative calculation $P_{\text{prt}}(r)$ and the actual $P(r)$. (a) is for the rotation and (b) is for the generic deformation, (c) is the same as (b), giving a zoom over the "birth" of the nonperturbative semiclassical saturation profile.

marked by the nonuniversal parametric scale δx_{NU} . For δx larger than δx_{NU} , the nonuniversal structure of the core is exposed.

What is the semiclassical structure of the LDOS? Time-domain semiclassical considerations [See Eq. (10) of Ref. [8]

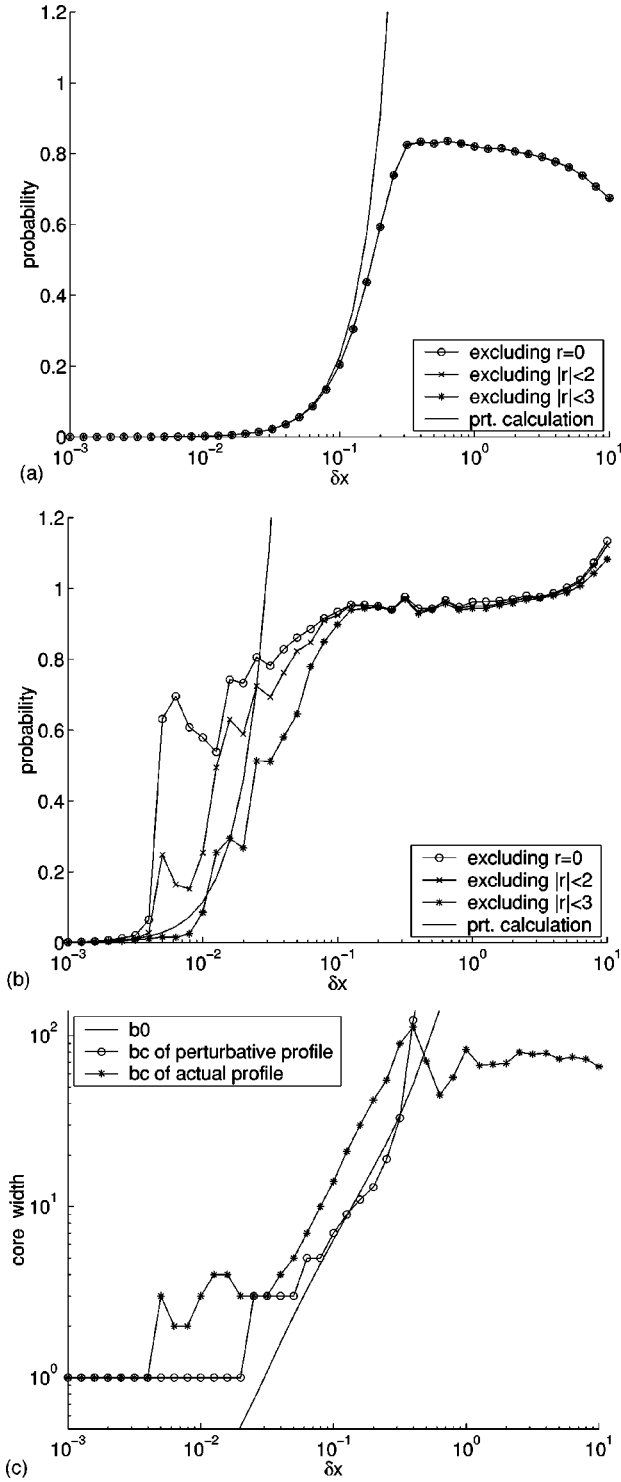


FIG. 8. (a) and (b) are $p(\delta x)$ for the rotation and for the generic deformation respectively. (c) is the perturbative calculation of the core width b_0 for the generic deformation. Also displayed is the calculated b_c for $P_{\text{pt}}(r)$, and the actual b_c for $P(r)$.

and related discussion there] imply that the nonuniversal structure of the core is

$$P(n-m) \approx \frac{1}{\pi} \frac{\delta E_{\text{SC}}}{\delta E_{\text{SC}}^2 + (E_n - E_m)^2} \quad (31)$$

with $\delta E_{\text{SC}} = \hbar / \tau_{\text{col}}$. The definition of the collision rate $1/\tau_{\text{col}}$ is similar to that of $1/\tau_{\text{bl}}$. The former is the collision rate with the deformed area A_{dfr} , and therefore it may be smaller than $1/\tau_{\text{bl}}$, because τ_{bl} is related to the total area A .

If, by mistake, we identified τ_{col} with τ_{bl} , then we would get the random wave result which is derived in the Appendix. This would be an over simplification. If it were true, it would imply that for $\delta x > \delta x_{\text{NU}}$ we should have got the same Lorentzian distribution in both cases (the rotation and the generic deformation). What we see, as a matter of fact, is that for the rotation (Fig. 5) we have a reasonably good agreement with Lorentzian whose width is $\delta k_{\text{SC}} = 1.5$, whereas for the generic deformation (Fig. 6) there is rough agreement with Lorentzian whose width is $\delta k_{\text{SC}} = 0.26$. The smaller width in the latter case clearly reflects having larger τ_{col} . For the generic deformation we also have pronounced non-Lorentzian features. Actually the global fitting to Lorentzian is quite bad. Our understanding is that these features are due to the bouncing-ball trapping: It leads to non-exponential decay of the time-dependent survival probability (see Ref. [8] for definition of the latter term), and hence to the observed non-Lorentzian features of the spreading profile.

We turn now to explain how δx_{NU} is determined. By definition it is the deformation which is required in order to expose features of the semiclassical landscape. These features start to be exposed once $\Gamma(\delta x) \sim \delta E_{\text{SC}}$ which should be converted into an equivalent expression $\delta x \sim \delta x_{\text{NU}}$. Thus we get

$$\delta x_{\text{NU}} = \left(\frac{A_{\text{dfr}}}{A_{\text{eff}}} \right)^{1/2} \times \lambda_B \sim \lambda_B. \quad (32)$$

Here A_{dfr} is the geometric area of the deformation (in the sense of scattering cross section), while A_{eff} is the effective area of the deformation. The definition of the latter is implied by comparing Eq. (27) with Eq. (25). By rescaling $D(s) \mapsto \alpha D(s)$ and $\delta x \mapsto \delta x / \alpha$, we can make $A_{\text{eff}} = A_{\text{dfr}}$ by convention. For special deformations δx_{NU} coincides with δx_c^{qm} implying that we get into the non-universal regime as soon as we have a breakdown of standard perturbation theory.

Our theoretical consideration so far do not imply a total collapse of perturbation theory. We may have in principle a coexistence of nonuniversal core component and perturbative tails. Actually we see such coexistence in the lower subfigures of Fig. 7, mainly for $\delta x = 0.2512$. The right peak around $\kappa = 0$ is clearly nonperturbative, while the rest of the profile is in reasonable (though not very good) agreement with the perturbative calculation. We are going to explain in the next section that the total collapse of perturbation theory for $\delta x > 0.3$ is actually not related at all to the appearance of nonuniversal features in the core structure. It is only circumstantial that in the hard wall limit this collapse happens as soon as we enter the nonuniversal regime.

XI. THE COLLAPSE OF PERTURBATION THEORY

A good starting for the following discussion is to consider the classical approximation for $P(r)$. Namely, in Eq. (1) one

approximates $\rho_n(Q,P)$ and $\rho_m(Q,P)$ by microcanonical distributions. A phase-space illustration of the energy surfaces which support $\rho_n(Q,P)$ and $\rho_m(Q,P)$ can be found in Fig. 1 of Ref. [8]. The classical $P(n|m)$ equals to the overlap of these surfaces.

If we were dealing with a generic system we could introduce a linearized version of the Hamiltonian $\mathcal{H}(Q,P;x) = \mathcal{H}_0(Q,P) + \delta x \mathcal{F}(Q,P; \delta x = 0)$, where $\mathcal{F}(Q,P;x) = \partial \mathcal{H} / \partial x$. By definition this linearization is a good approximation provided $\delta x \ll \delta x_c^{\text{cl}}$, where δx_c^{cl} is the classical correlation scale of $\mathcal{F}(Q,P;x)$ with respect to x . In the classical linear regime the classical $P(r)$ has the scaling property $P(r) = 1/\delta x \hat{P}(r/\delta x)$. See Ref. [15] for details.

An equivalent definition of δx_c^{cl} in the quantum-mechanical case is obtained by looking on the x dependence of the matrix elements of $\mathcal{F}(Q,P;x)$ in some *fixed basis*. Again we define δx_c^{cl} as the respective correlation scale. It is quite clear that for cavity with soft walls we have

$$\delta x_c^{\text{cl}} = l \quad [\text{soft walls}], \quad (33)$$

where $l = E/f$ has been defined as the penetration distance upon collision. From a purely classical point of view the hard wall limit $l \rightarrow 0$ is a *nonlinear* limit. But this is not true quantum mechanically. Here we have

$$\delta x_c^{\text{cl}} \sim \lambda_B \quad [\text{hard walls}] \quad (34)$$

for $l < \lambda_B$. The terminology ‘‘classical correlation scale’’ while referring to δx_c^{cl} becomes misleading here, but we shall keep using it anyway.

The theory of the core-tail structure [14] is valid only in the linear regime $\delta x \ll \delta x_c^{\text{cl}}$. Let us assume for a moment soft walls. We can ask is $\delta x \ll \delta x_c^{\text{cl}}$ a sufficient condition for having a core-tail structure? The answer is definitely not. Perturbation theory has a final collapse once the core width $b_0(\delta x)$ becomes of the order of the band width b . This defines a parametric scale δx_{prt} . For $\delta x > \delta x_{\text{prt}}$ the LDOS becomes purely nonperturbative. In Wigner’s theory of random banded matrices this corresponds to the crossover from Lorentzian to semicircle line shape.

However, in the limit of hard walls the above mechanism of collapse becomes irrelevant because the band width is infinite ($b = \infty$). On the other hand we still have to satisfy the inequality $\delta x \ll \delta x_c^{\text{cl}}$. Thus, for $\delta x > \lambda_B$ we expect a total collapse of perturbation theory, as indeed observed in the numerical study.

ACKNOWLEDGMENTS

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APPENDIX: OVERLAP OF UNCORRELATED RANDOM WAVES

It is possible to estimate the overlap $|\langle n|m \rangle|^2$ if we assume that $|n\rangle$ and $|m\rangle$ are *uncorrelated* random superpositions of plane waves: A random superposition of plane

waves is characterized by the correlation function

$$\langle \psi_R(\mathbf{x}_1) \psi_R(\mathbf{x}_2) \rangle = \frac{1}{\mathcal{V}} \cos(k|\mathbf{x}_2 - \mathbf{x}_1|),$$

where $\cos(kr) \equiv \langle \exp(i\mathbf{k} \cdot \mathbf{r}) \rangle_\Omega$ is a generalized Bessel function (for further details see Appendix D of Ref. [14]). Assuming that the wave function of $|n\rangle$ is uncorrelated with the wave function of $|m\rangle$ one obtains

$$\begin{aligned} \langle |\langle n|m \rangle|^2 \rangle &= \left(\frac{1}{\mathcal{V}} \right)^2 \int \int \cos(k_n |\mathbf{x}_2 - \mathbf{x}_1|) \\ &\quad \times \cos(k_m |\mathbf{x}_2 - \mathbf{x}_1|) d\mathbf{x}_1 d\mathbf{x}_2. \end{aligned}$$

The integration is over the whole volume of the cavity. Using the definition of the cos function we can cast this expression into the form $\langle |\langle n|m \rangle|^2 \rangle = \langle f(\mathbf{q}) \rangle_q$ where the average is over the difference $\mathbf{q} = (\mathbf{k}_2 - \mathbf{k}_1)$, with all possible orientations for $|\mathbf{k}_1| = k_m$ and for $|\mathbf{k}_2| = k_n$. The function $f(\mathbf{q})$ is defined as

$$f(\mathbf{q}) = \left| \int e^{i\mathbf{q} \cdot \mathbf{x}} d\mathbf{x} \right|^2.$$

The function $f(\mathbf{q})$ depends mainly on $q = |\mathbf{q}|$. We can obtain an estimate for $f(q)$ by considering a spherical cavity $|\mathbf{x}| < R$ with the same volume. Using spherical coordinates it is straightforward to obtain

$$f(q) \approx \left(\text{FT} \left[\frac{1}{d-1} (R^2 - x^2)^{(d-1)/2} \right] \right)^2,$$

where FT is a Fourier transform from x to q . For $q \ll 1/R$ we have simply $f(q) = (\mathcal{V})^2$. For $q \gg 1/R$ we have $f(q) = (\mathcal{V})^2 / (Rq)^{d+1}$. This is because of the singularity at $x = \pm R$. The average over q can be done using again spherical coordinate. We have $q \approx [(\Delta k)^2 + 2k^2(1 - \cos \theta)]^{1/2}$, where $\Delta k = |k_n - k_m|$, and we can transform the $d\theta$ integration into a dq integration

$$\begin{aligned} \langle f(\mathbf{q}) \rangle_q &= \frac{\Omega_{d-2}}{\Omega_{d-1}} \\ &\quad \times \int \frac{qdq}{k^2} \left(\frac{\sqrt{[q^2 - (\Delta k)^2][(2k)^2 - q^2]}}{2k^2} \right)^{d-3} f(q), \end{aligned}$$

where Ω_d is the solid angle in d dimensions. There is no point in trying to carry an exact integration. Rather, it is important to observe that a practical approximation for the overlap is

$$\langle |\langle n|m \rangle|^2 \rangle \approx \left(\frac{1}{kR} \right)^{d-1} \frac{1}{1 + [R \cdot (k_n - k_m)]^2}.$$

In the latter expression we have neglected the d -dependent normalization prefactor. It is a ‘‘practical’’ approximation since it gives the correct behavior for both small and large values of Δk . The interpolation around $\Delta k \sim 1/R$ cannot be trusted.

- [1] The classical approximation for $P(n|m)$ works well for ergodic eigenstates. As explained later in the text we actually consider the averaged profile. Namely, $P(n|m)$ is regarded as a function of $r=n-m$, and it is averaged over the reference state m . With this procedure the effect of the minority of non-ergodic eigenstates can be neglected in the classical limit.
- [2] The two parametric scales of Wigner's RMT model correspond to δx_c^{qm} and δx_{prt} of Table I. Consequently there are three parametric regimes in Wigner's theory. These are the standard perturbative regime ($\delta x \ll \delta x_c^{\text{qm}}$), the Lorentzian regime ($\delta x_c^{\text{qm}} \ll \delta x \ll \delta x_{\text{prt}}$), and the semi-circle regime ($\delta x \gg \delta x_{\text{prt}}$). As an artifact of this RMT model, there actually exists a fourth regime (Anderson strong localization regime) that in the present context does not have a semiclassical analog.
- [3] As explained in Ref. [8] the generic hierarchy is $\delta x_{\text{NU}} \ll \delta x_{\text{prt}} \ll \delta x_{\text{SC}} \ll \delta x_c^{\text{cl}}$. This hierarchy is realized in the classical limit (small \hbar) where we have soft walls (see Sec. II). The hard wall limit is nongeneric, and these four parametric scales coincide. In the present paper we assume hard walls unless stated otherwise.
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- [21] In order to find the average profile we divide the κ axis into small bins. We find the average $P(n|m)$ value in each bin. In order to get $P(r)$ we resample the average profile. The distance between the sampling points is equal to the mean level spacing. We are not interested in features on scale of the mean level spacing. Therefore the transformation from the continuous variable κ to the integer variable r should not be considered as problematic.