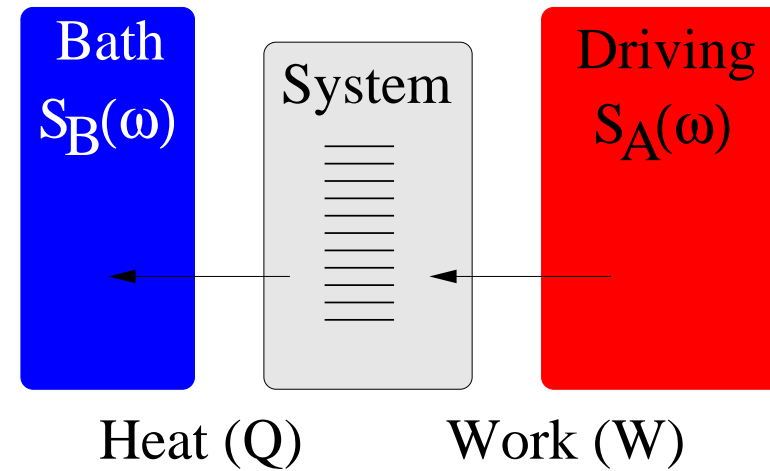
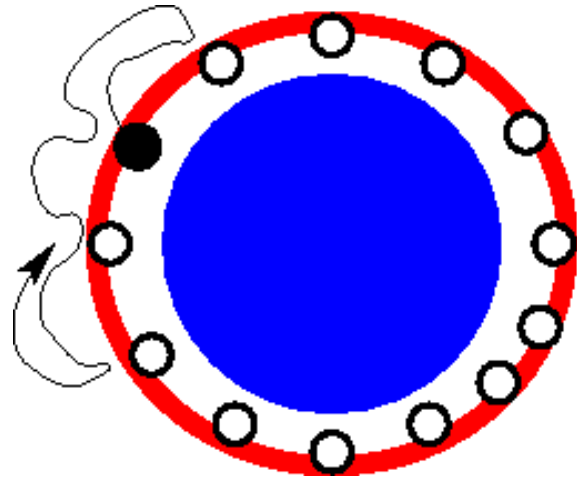


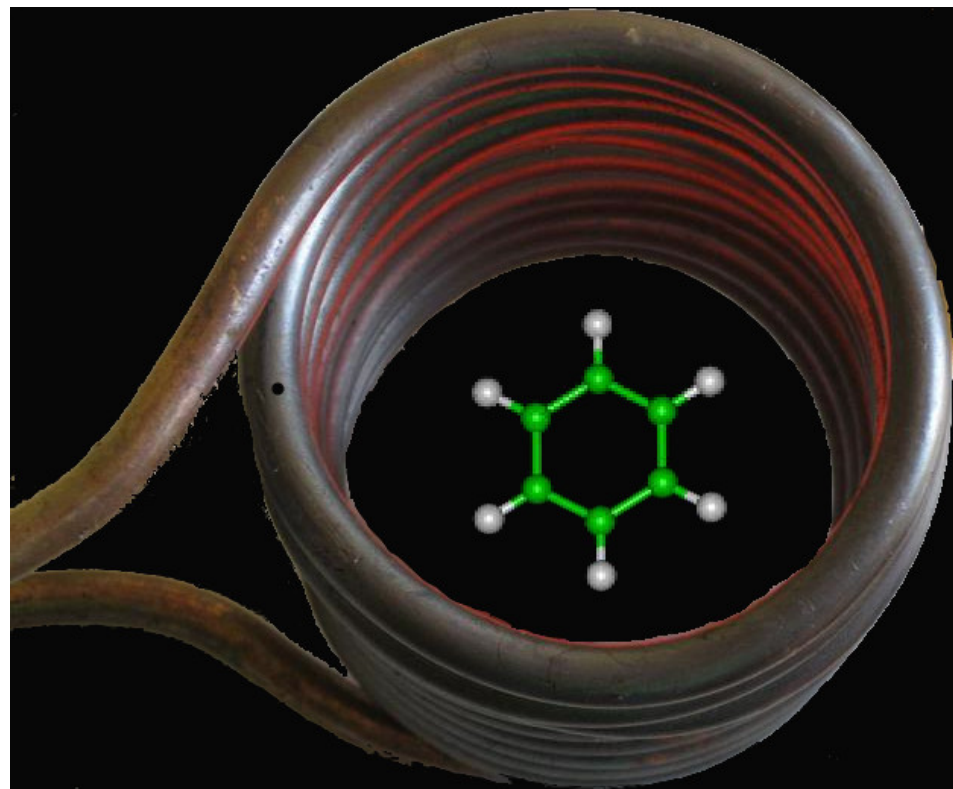
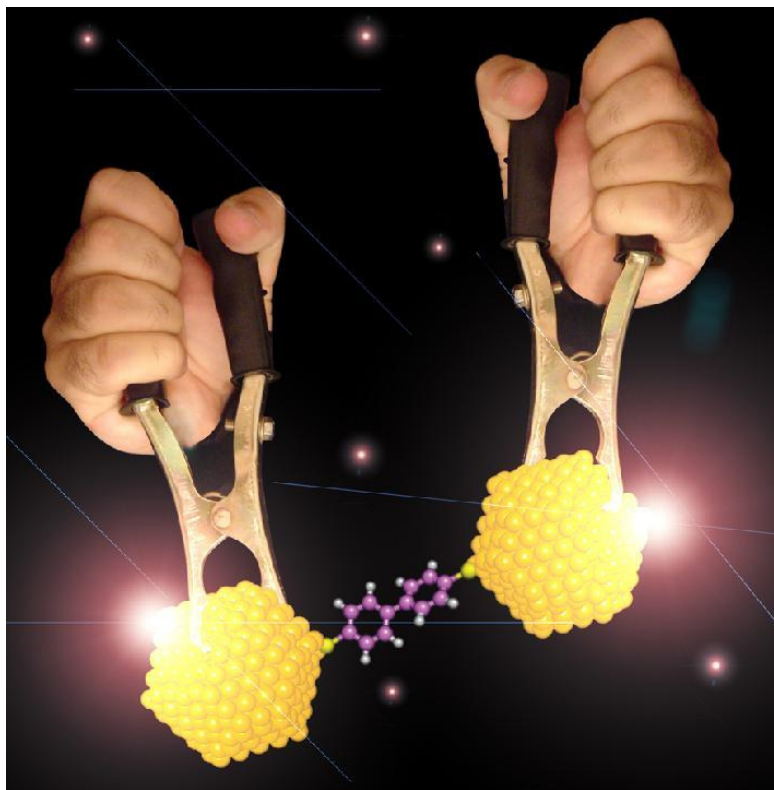
# The non-equilibrium steady state of sparse systems with non trivial topology

Daniel Hurowitz, Ben-Gurion University



- [1] D. Hurowitz and D. Cohen, Europhysics Letters 93, 60002 (2011)
- [2] D. Hurowitz, S. Rahav and D. Cohen, Europhysics Letters 98, 20002 (2012)
- [3] D. Hurowitz, S. Rahav and D. Cohen, arXiv (2013)

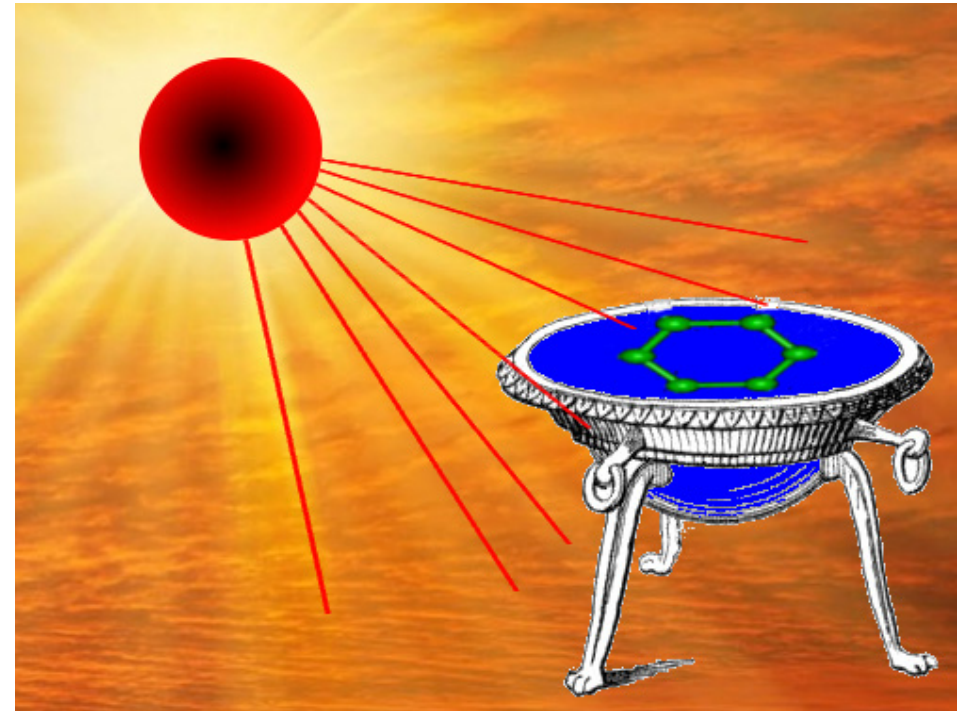
## Let us go wireless



We would like to induce current in a closed device (no leads),  
even if the the particles have no charge.

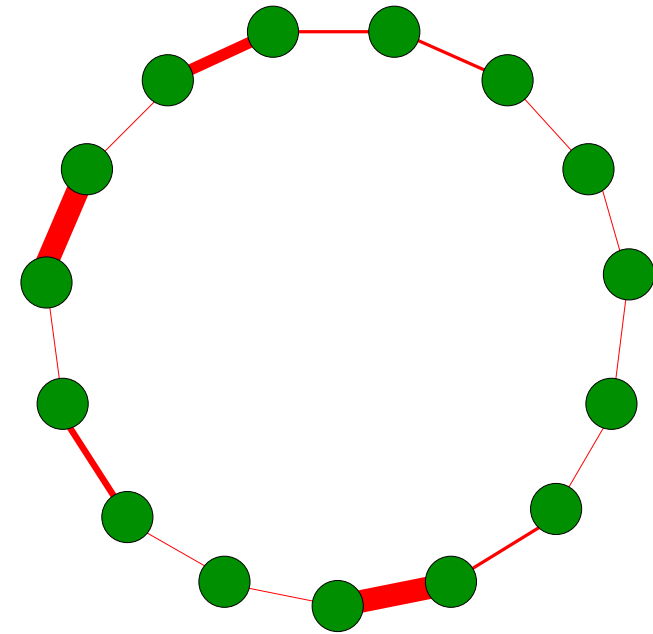
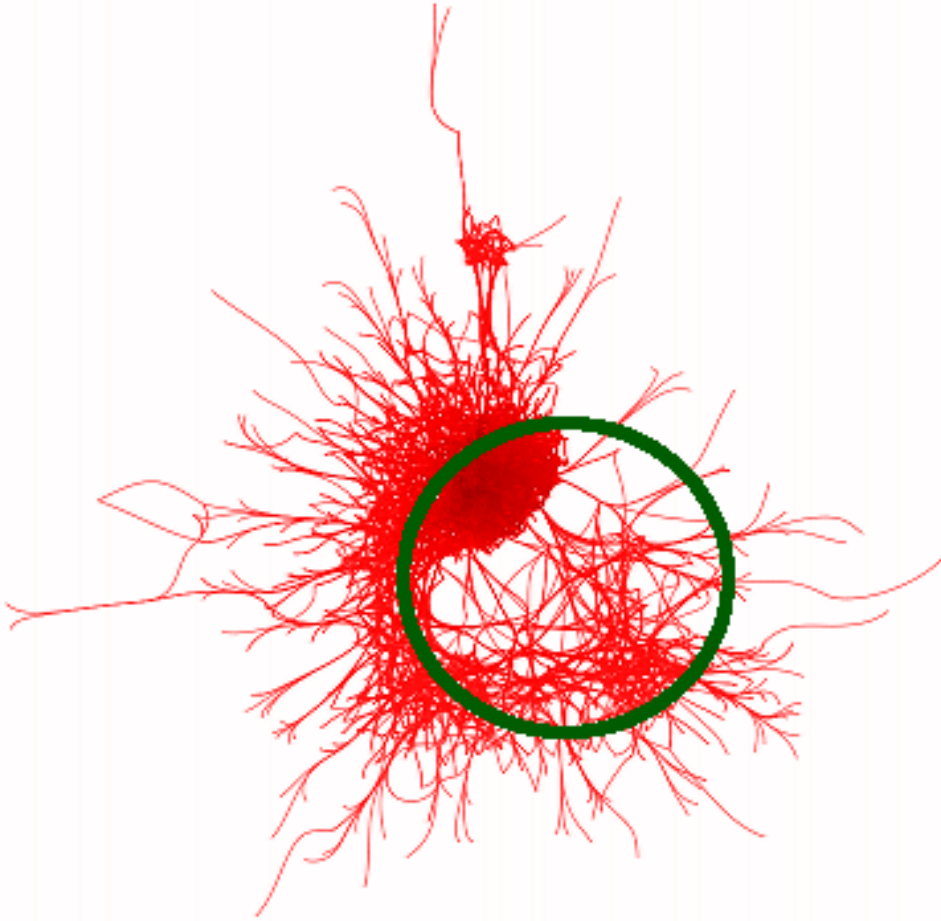
[Left figure is courtesy of Amir Yacobi]

# (Non Equilibrium) Statistical Mechanics of Small Systems



We would like to push down the laws of thermodynamics  
into the mesoscopic scale,  
where fluctuations and quantum mechanics dominate.

## Sparse systems



In our study we consider systems that are "sparse" or "glassy", meaning that many time scales are involved.

Standard thermodynamics does not apply to such systems.

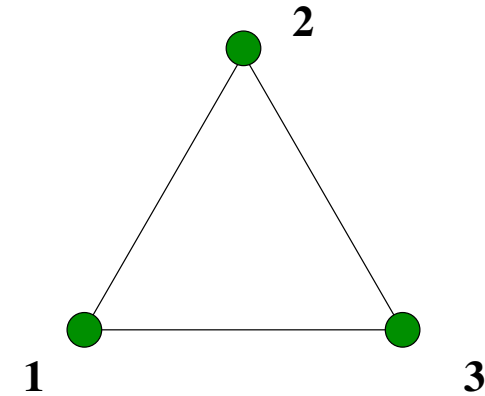
# Non Equilibrium Statistical Mechanics

Master Equation:  $\dot{\mathbf{p}} = \mathcal{W}\mathbf{p}$

$$\dot{p}_1 = -(w_{21} + w_{13})p_1 + w_{12}p_2 + w_{13}p_3$$

$$\dot{p}_2 = w_{21}p_1 - (w_{12} + w_{32})p_2 + w_{23}p_3$$

$$\dot{p}_3 = w_{31}p_1 + w_{32}p_2 - (w_{13} + w_{23})p_3$$



Stochastic fields:

$$\mathcal{E}_{12} = \ln \frac{w_{12}}{w_{21}}$$

$$\mathcal{E}_{23} = \ln \frac{w_{23}}{w_{32}}$$

$$\mathcal{E}_{31} = \ln \frac{w_{31}}{w_{13}}$$

In equilibrium

$$\mathcal{E}_{12} + \mathcal{E}_{23} + \mathcal{E}_{31} = 0$$

$$\mathcal{E}_{12} = E_2 - E_1$$

$$\mathcal{E}_{23} = E_3 - E_2$$

$$\mathcal{E}_{31} = E_1 - E_3$$

$$p_i \propto \exp(-E_i/T)$$

Non equilibrium

$$\mathcal{E}_{12} + \mathcal{E}_{23} + \mathcal{E}_{31} \neq 0$$



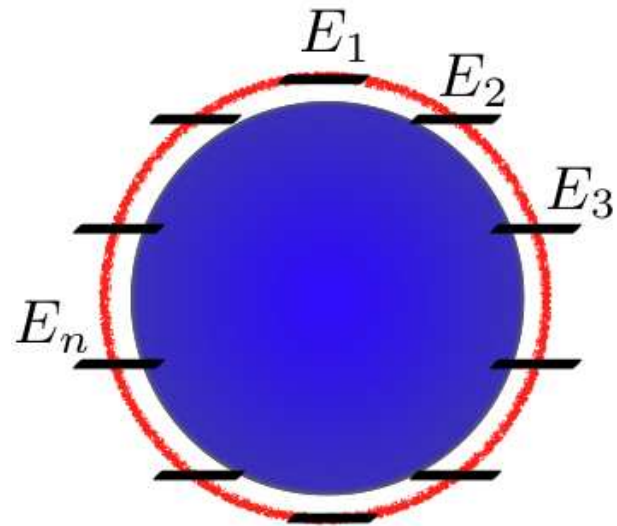
$$w_{12}w_{23}w_{31} \neq w_{13}w_{32}w_{21}$$

If  $\oint \mathcal{E}(x)dx = 0$  for all closed loops the steady state will be an equilibrium state.

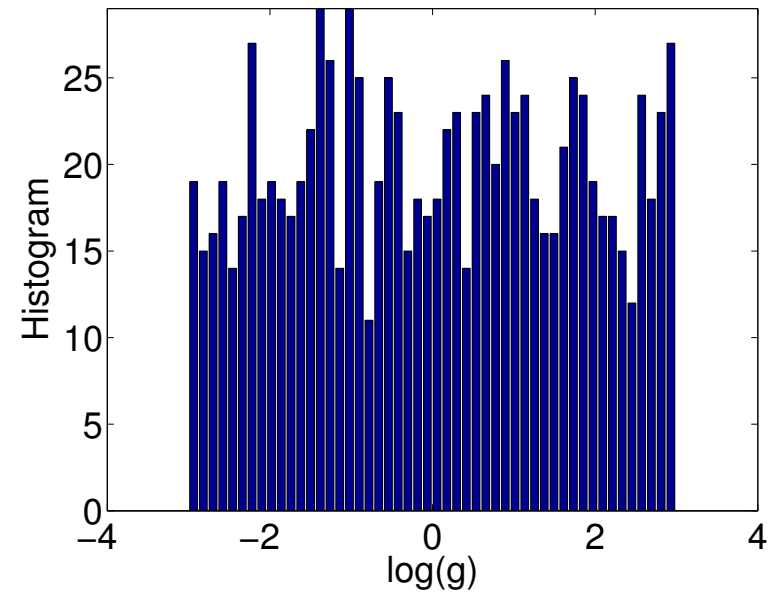
Otherwise, the system will reach a **Non-Equilibrium Steady State (NESS)**.

# The model

System + Bath + Driving



Histogram of couplings



$$w_{n+1,n}^{total} = w_{n+1,n}^{\beta} + \nu g_n$$

$w^{\beta}$  corresponds to  $T_B = \text{finite}$

$w^{\nu}$  corresponds to  $T_A = \infty$

← few decades →

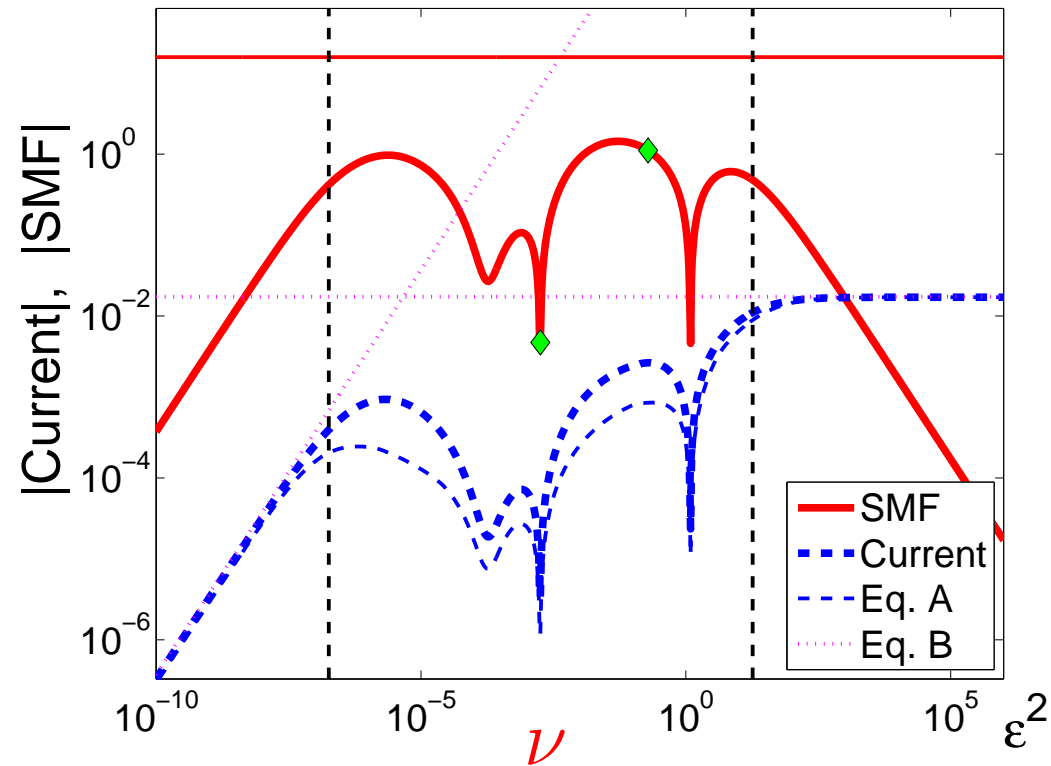
$g_n = \text{couplings}$

“sparsity” = log wide distribution of couplings

## Current vs. driving

Driving  $\rightsquigarrow$  Stochastic Motive Force  $\rightsquigarrow$  Current

Regimes: LRT regime, Sinai regime, Saturation regime



Due to the **sparsity**, we have an intermediate **Sinai regime**.

The width of the Sinai regime is determined by the **log-width of the distribution**.

# Sinai Diffusion

Conventional random walk :

Equal & symmetric transition rates

$$I \propto \frac{1}{N}$$

Sinai Random Walk:

Uncorrelated & non symmetric transition rates

↪ build up of activation barrier

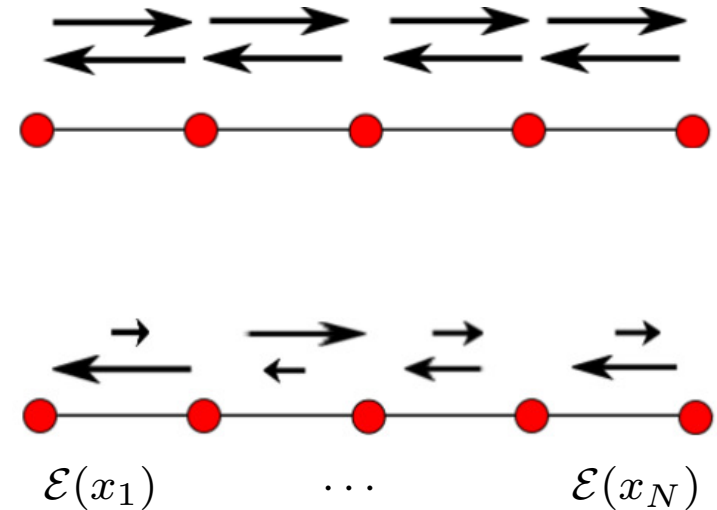
$$I \propto e^{-\sqrt{N}}$$

Our model:

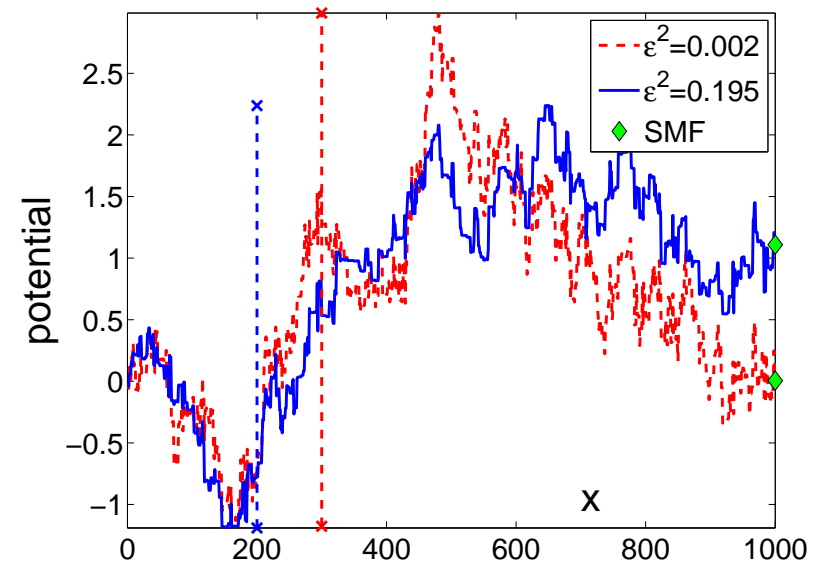
Telescopic correlations:  $\mathcal{E}(x_n) \sim \Delta_n \equiv (E_n - E_{n+1})$

Yet... we have sparsely distributed couplings

$$I \sim \frac{1}{N} \bar{w} \exp \left[ -\frac{\mathcal{E}_n}{2} \right] 2 \sinh \left( \frac{\mathcal{E}_0}{2} \right)$$

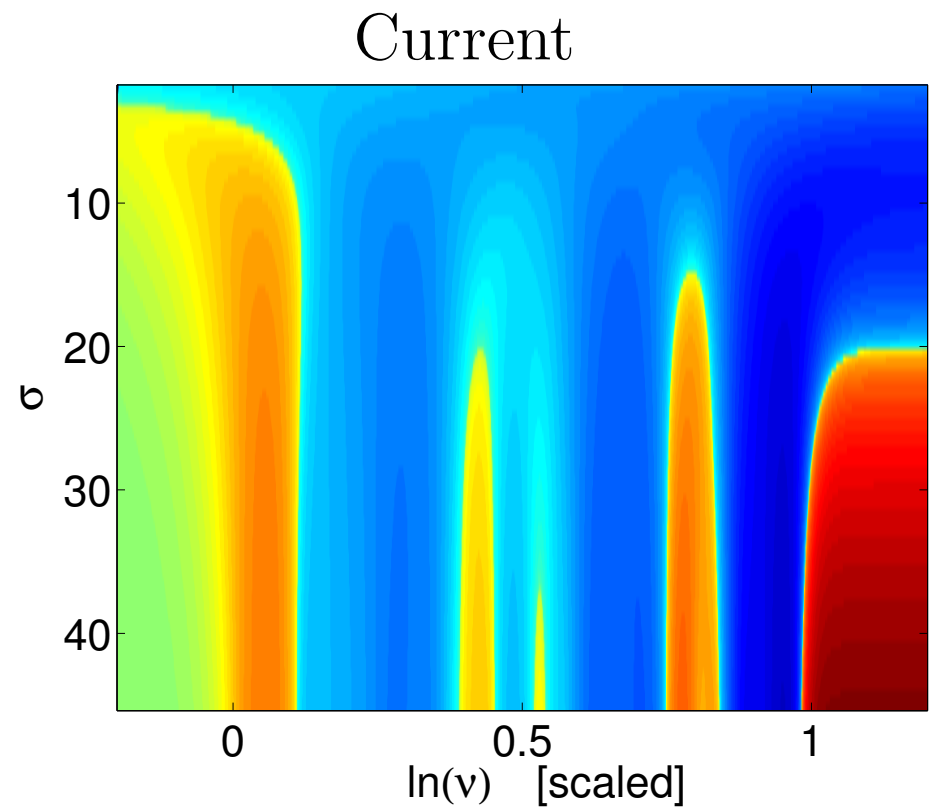
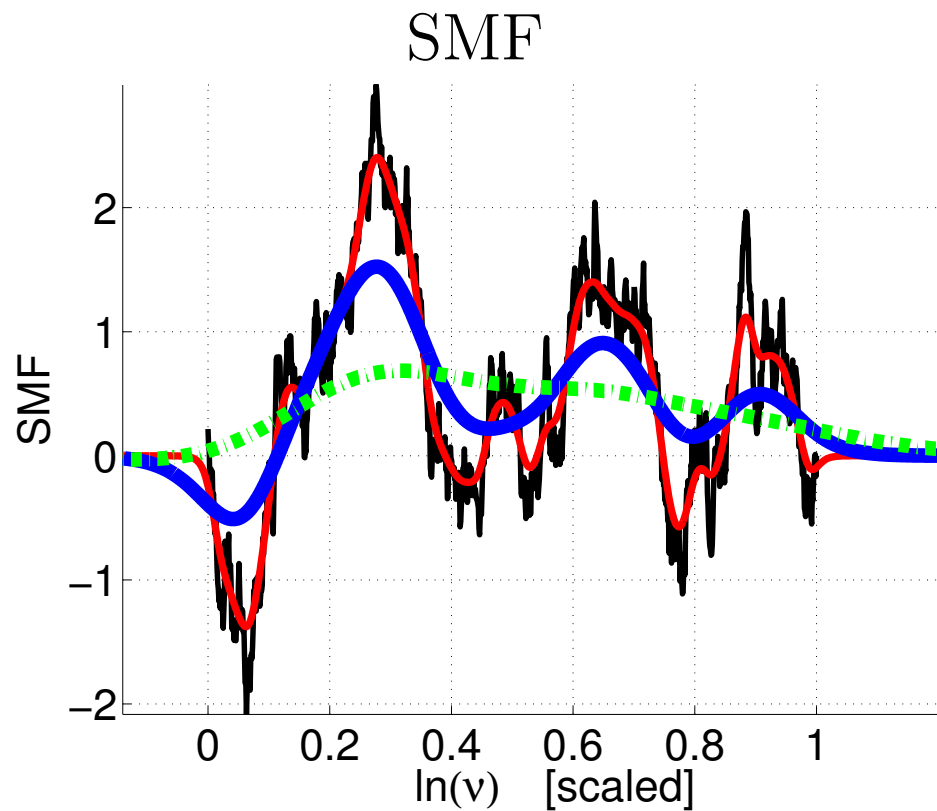


$$\int_0^x \mathcal{E}(x') dx' \sim \sqrt{N}$$





## Direction of the Current in Sinai Regime



**The direction of the current is determined by the SMF**

## The SMF in the Sinai Regime

Stochastic field :

$$\mathcal{E}(x_n) \equiv \ln \left[ \frac{w_{\vec{n}}}{w_{\overleftarrow{n}}} \right] \approx - \left[ \frac{1}{1 + g_n \nu} \right] \frac{E_n - E_{n-1}}{T_B}$$

Stochastic Motive Force:

$$\mathcal{E}_{\circlearrowleft} \equiv \ln \left[ \frac{\prod_n w_{\vec{n}}}{\prod_n w_{\overleftarrow{n}}} \right] = \oint \mathcal{E}(x) dx$$

Coarse grained random walk:

$$\mathcal{E}_{\circlearrowleft}(\tau) = - \sum_{n=1}^N f_{\sigma}(\tau - \tau_n) \frac{E_n - E_{n-1}}{T_B}$$

$$\sigma = \ln \frac{g_{\max}}{g_{\min}}, \quad [\text{log-width of distribution}]$$

$$\tau \equiv \frac{1}{\sigma} \ln(g_{\max} \nu), \quad \tau_n = \frac{1}{\sigma} \ln \left( \frac{g_{\max}}{g_n} \right)$$

$$f_{\sigma}(t) \equiv [1 + e^{\sigma t}]^{-1} \quad [”step” function]$$

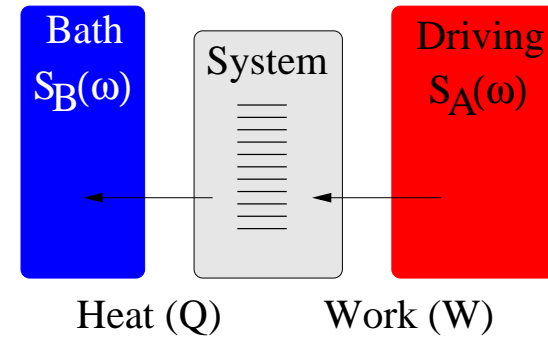
Expected number of sign changes  $\approx \sqrt{\pi \sigma}$

## FD phenomenology for a “sparse” system

$$W_{nm} = w_{nm}^\beta + \nu g_{nm}$$

$$\dot{W} = \text{rate of heating} = \frac{D(\nu)}{T_{\text{system}}}$$

$$\dot{Q} = \text{rate of cooling} = \frac{D_B}{T_B} - \frac{D_B}{T_{\text{system}}}$$



Hence at the NESS:

$$T_{\text{system}} = \left(1 + \frac{D(\nu)}{D_B}\right) T_B$$

$$\dot{Q} = \dot{W} = \frac{1/T_B}{D_B^{-1} + D(\nu)^{-1}}$$

Experimental way to extract response:

$$D(\nu) = \frac{\dot{Q}(\nu)}{\dot{Q}(\infty) - \dot{Q}(\nu)} D_B$$

$D(\nu)$  exhibits LRT to SLRT crossover

$$D(\nu) = \left[ \left( \frac{w_n}{w_\beta + w_n} \right) \right] \left[ \left( \frac{1}{w_\beta + w_n} \right) \right]^{-1}$$

$$D_{[\text{LRT}]} = \overline{g_n} \nu \quad [\text{weak driving}]$$

$$D_{[\text{SLRT}]} = \left[ \overline{1/g_n} \right]^{-1} \nu \quad [\text{strong driving}]$$

Expressions above assume n.n. transitions only.

## Summary of main results

1. Due to the **sparsity** of the perturbation matrix, the NESS is of glassy nature [1]
2. An extension of the **Fluctuation-Dissipation phenomenology** has been proposed [1]
3. A log-wide distribution of couplings leads to **Sinai-type physics** [2]
4. In the Sinai regime, the fluctuations of the current reflect the log-wide distribution of the couplings [3]